

Schwarzschild–Couder 型の CTA 小・中口径望遠鏡の開発状況

奥村 暁^{1,2}

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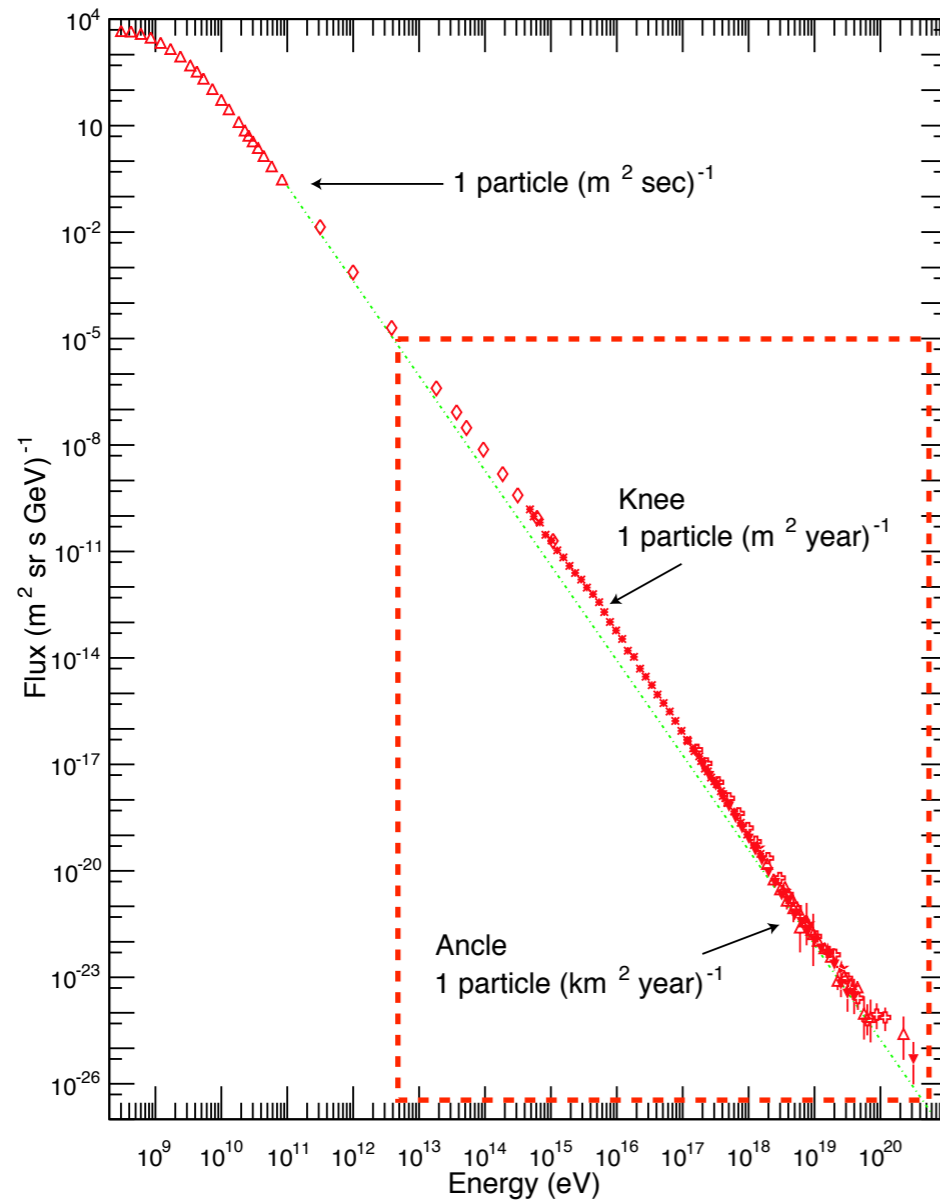
² 名古屋大学 素粒子宇宙起源研究所 (KMI)

³ 茨城大学 理学部

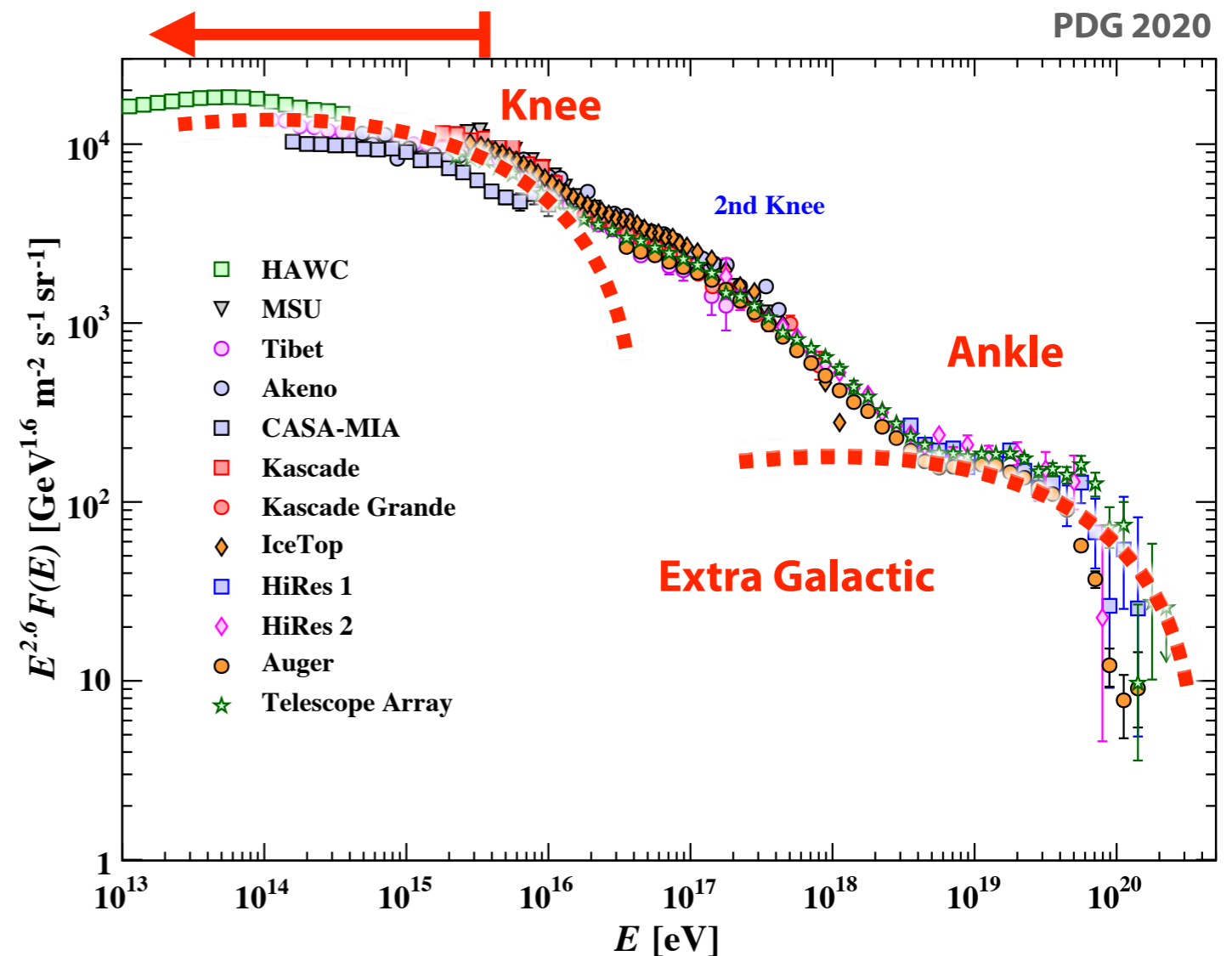
2020 年 9 月 9 日 @ 日本天文学会 2020 年秋季年会

Cosmic-ray Hadron Spectrum

Modified Cronin+ (1997)



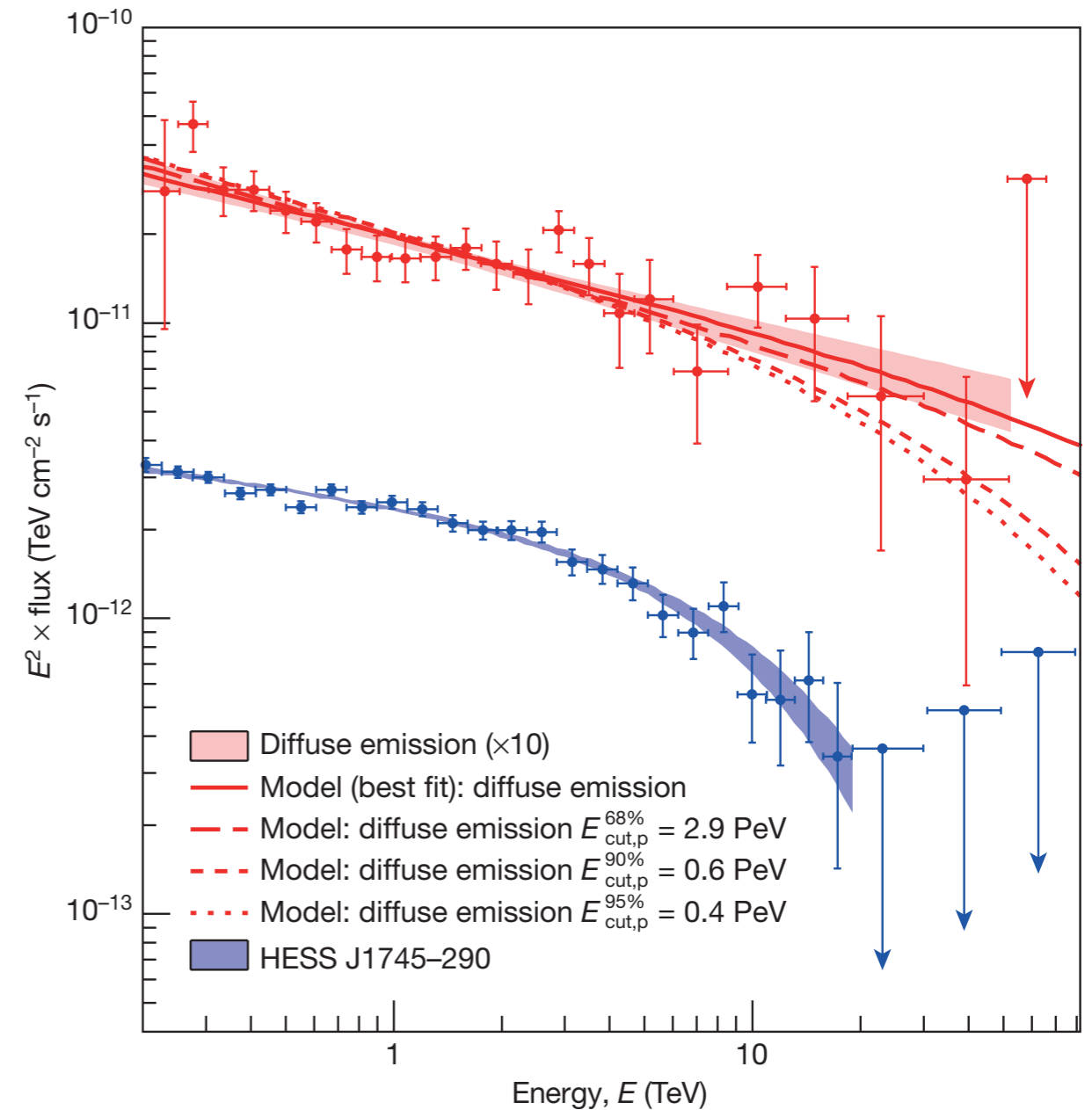
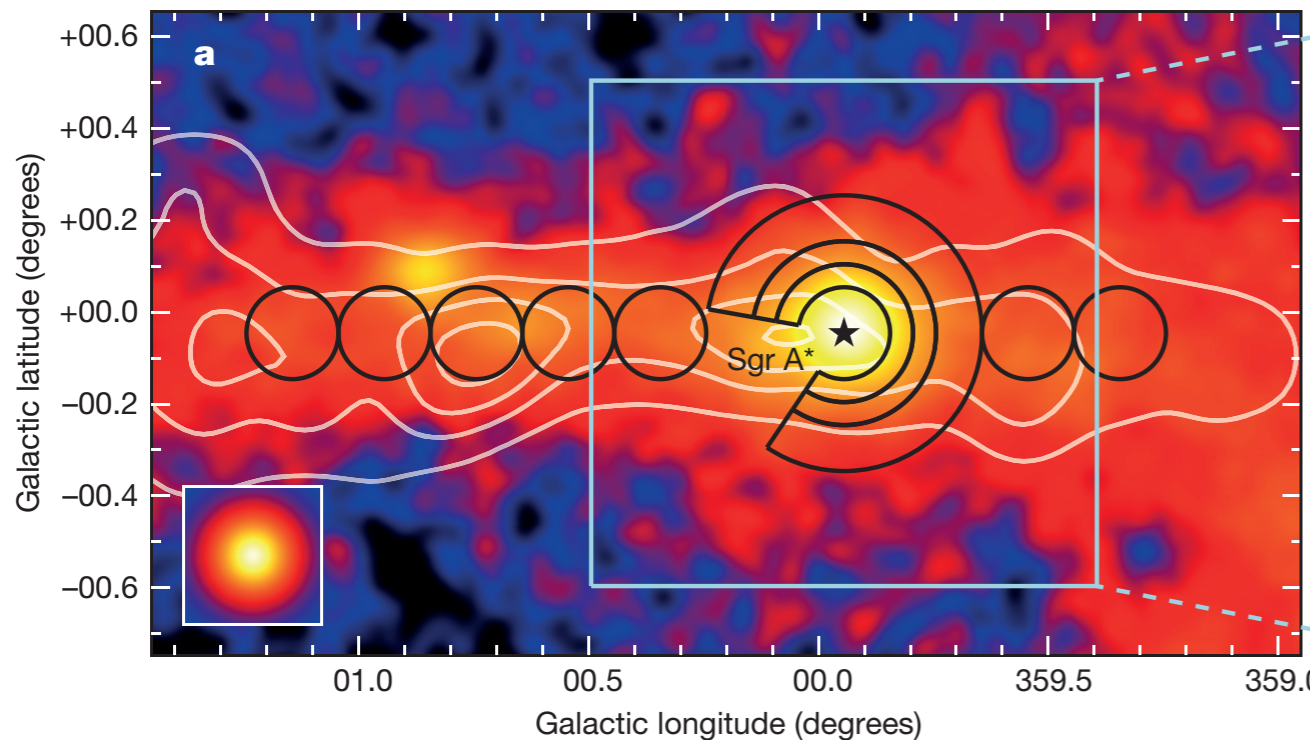
Galactic CRs $\sim 3 \times 10^{15}$ eV (3 PeV)



- $\sim 10^8$ eV (~ 100 MeV) to $> 10^{20}$ eV, with a power law of $dN/dE = E^{-2.7}$ to $E^{-3.0}$
- What is the origin (PeVatron) of Galactic CRs ($< \sim 3$ PeV)? Supernova remnants? Galactic center?

Galactic Center Region

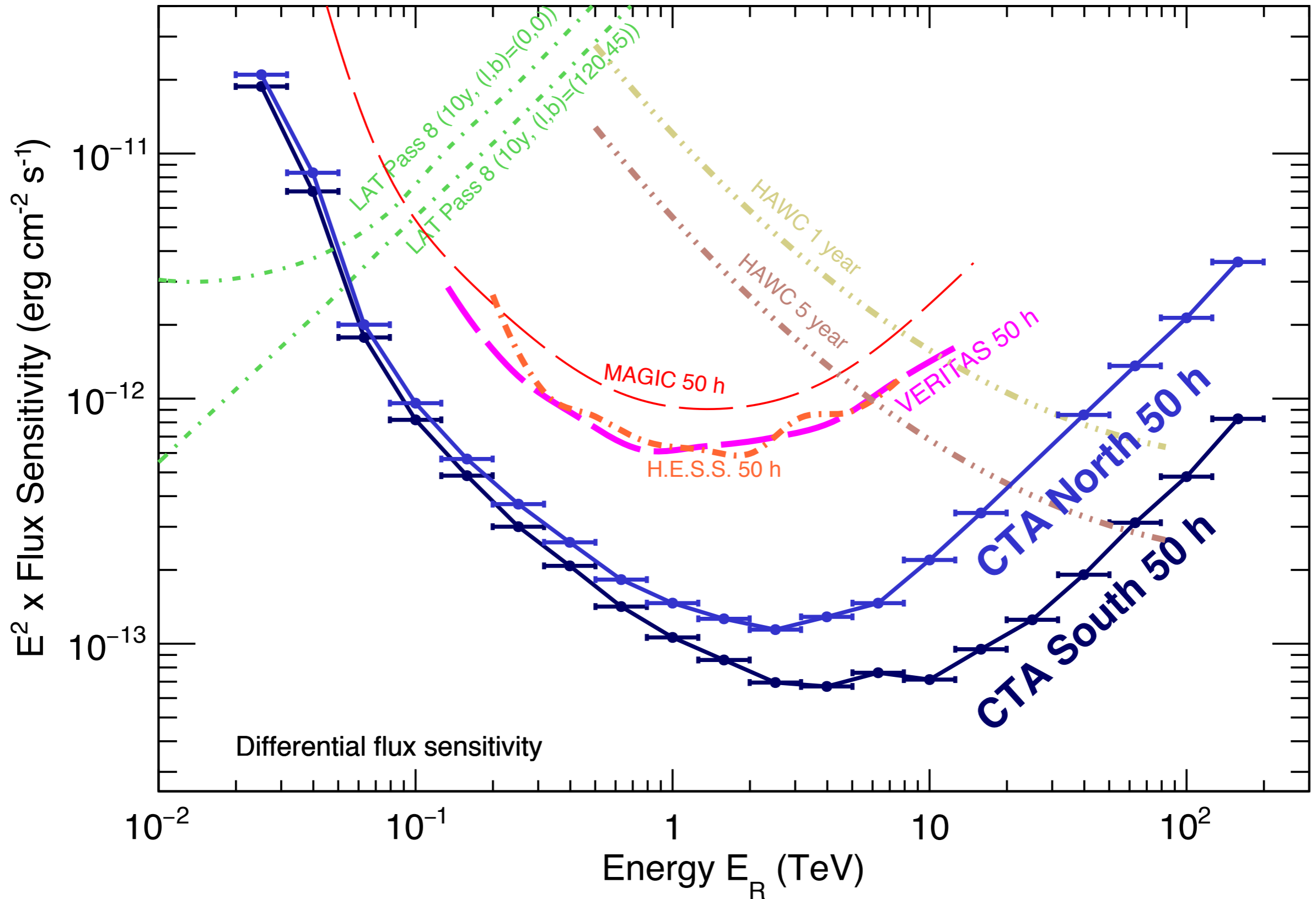
H.E.S.S. (2016)



- Massive black hole ($4 \times 10^6 M_{\odot}$) at Sgr A* (bright radio source)
- Point source HESS J1745-290 at Str A* and diffuse gamma-ray emission
- Diffuse component has a cutoff energy of 2.9 PeV (68% conf.) → PeVatron?

High-energy Frontier by CTA SSTs

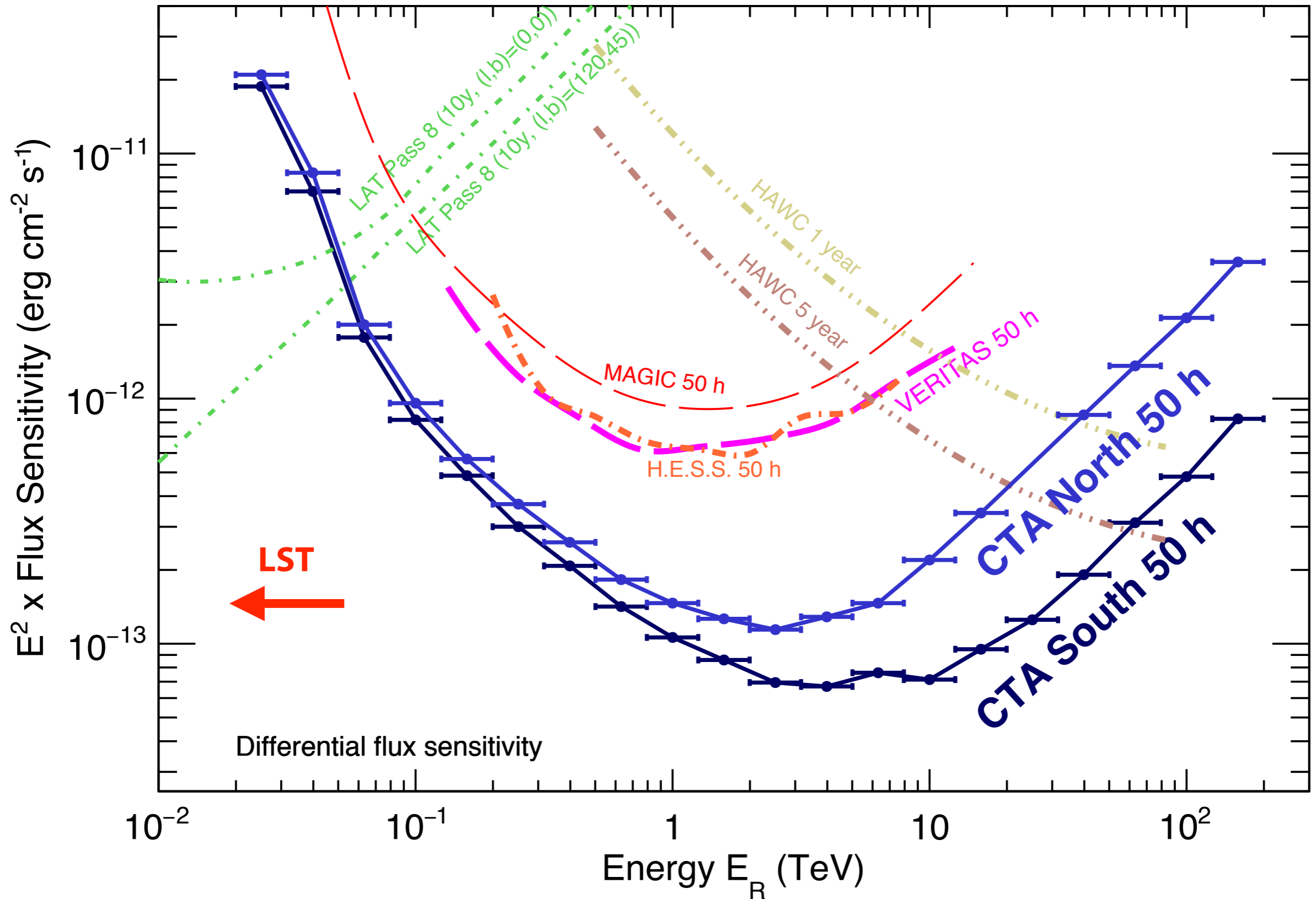
CTA Consortium arXiv:1709.07997



www.cta-observatory.org/science/cta-performance/ (prod3b-v1)

High-energy Frontier by CTA SSTs

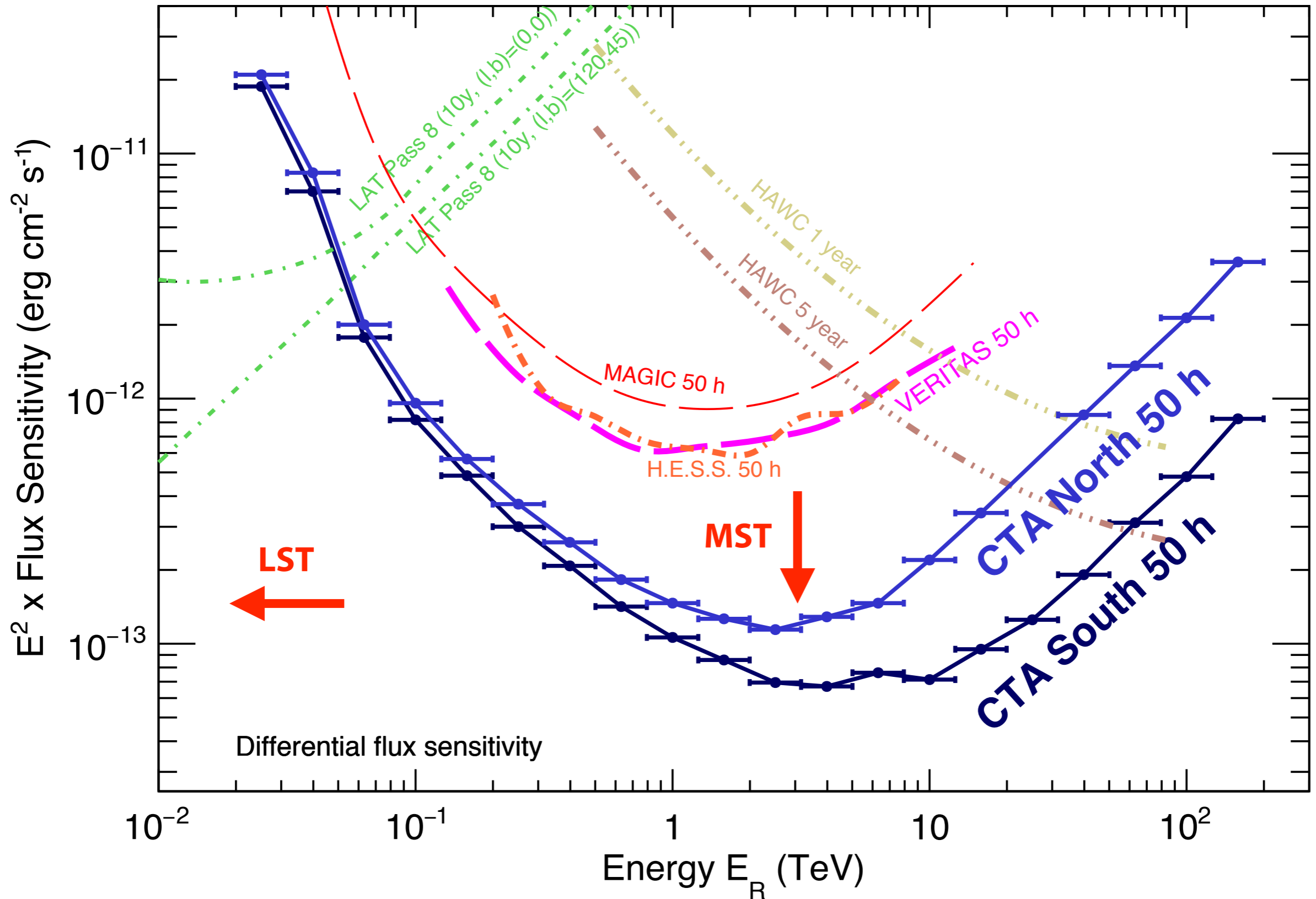
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High-energy Frontier by CTA SSTs

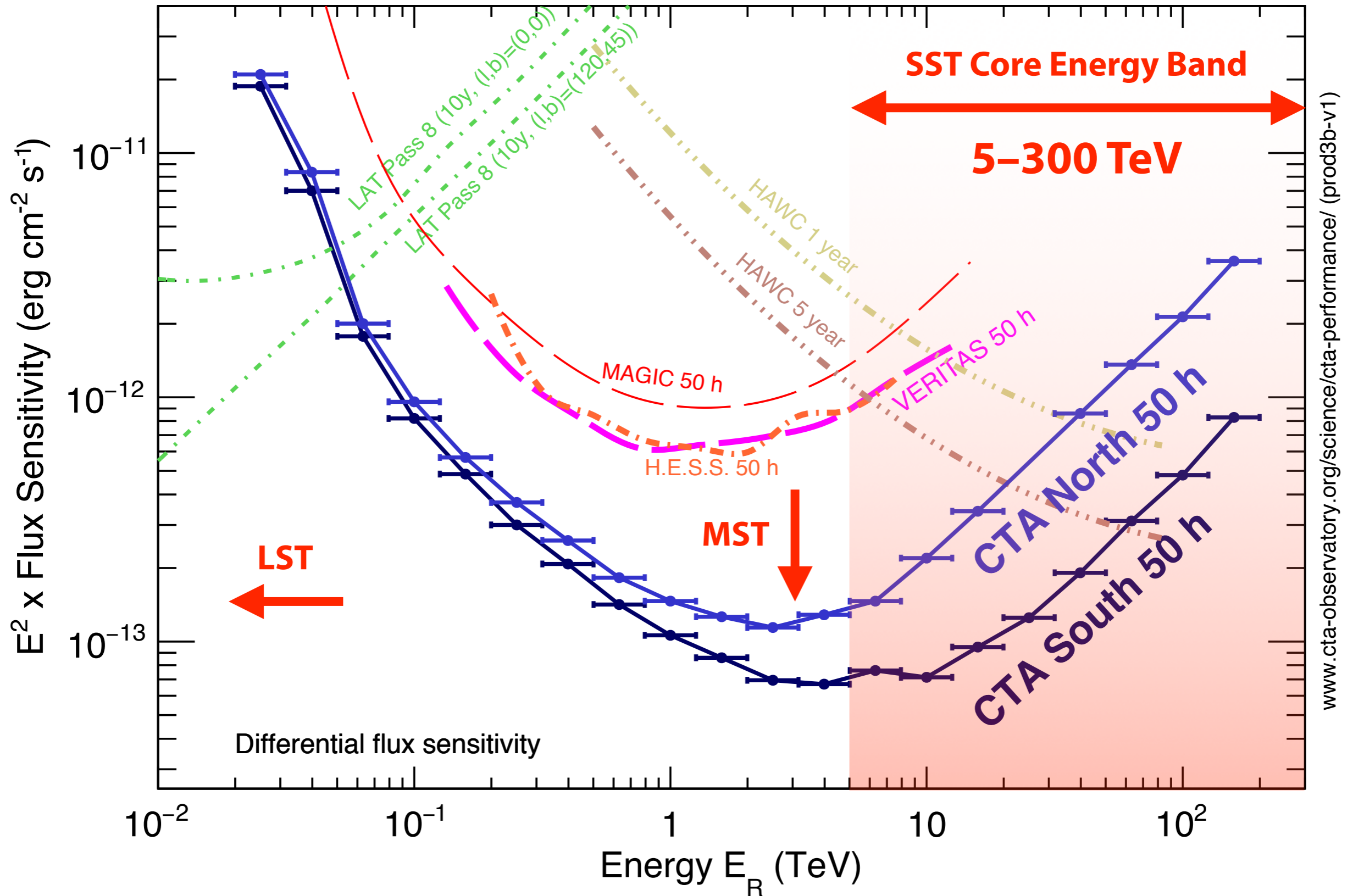
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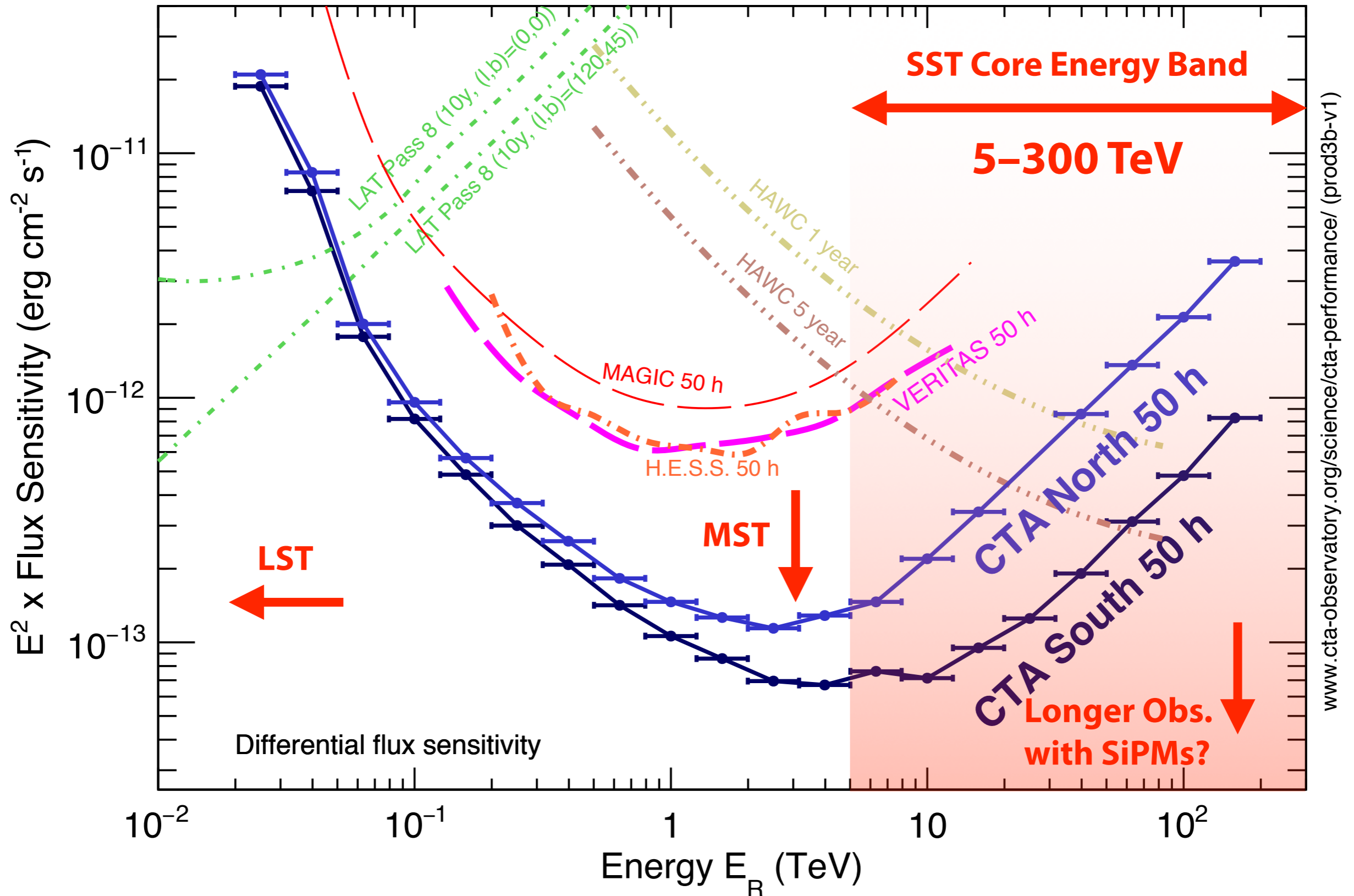
High-energy Frontier by CTA SSTs

CTA Consortium arXiv:1709.07997



High-energy Frontier by CTA SSTs

CTA Consortium arXiv:1709.07997



Cherenkov Telescope Array (CTA)



cherenkov
telescope
array

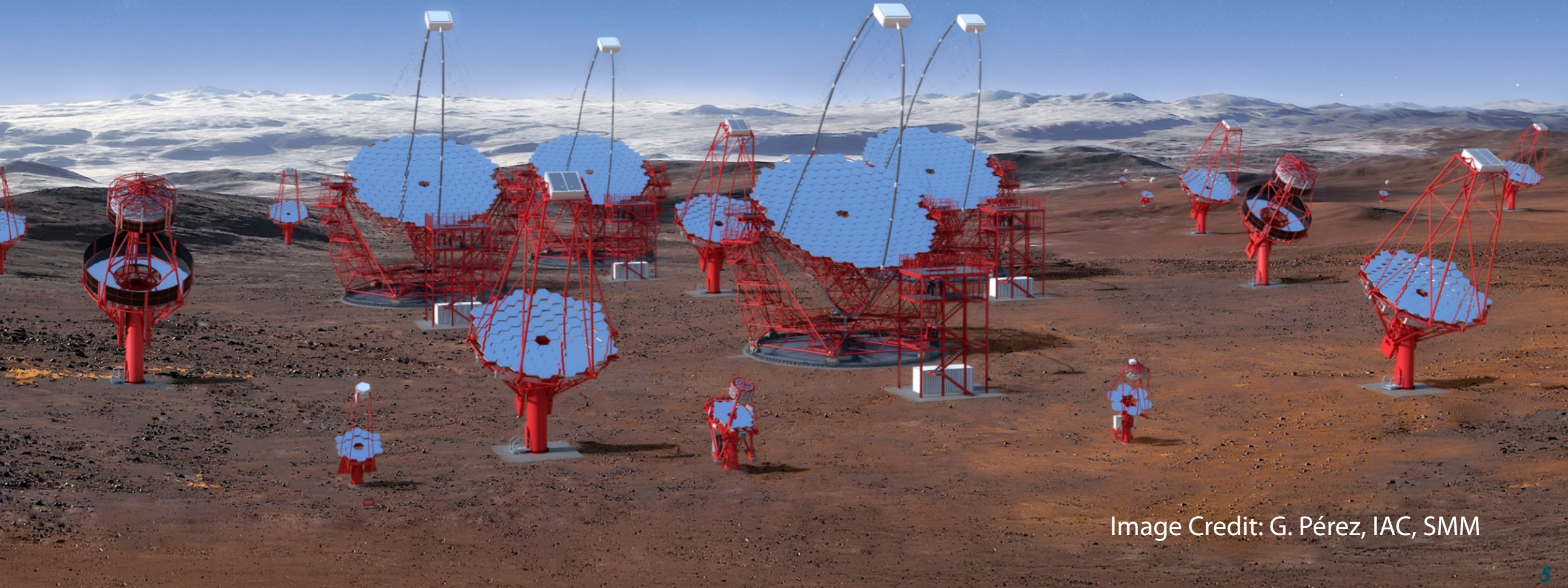


Image Credit: G. Pérez, IAC, SMM

Cherenkov Telescope Array (CTA)



cherenkov
telescope
array

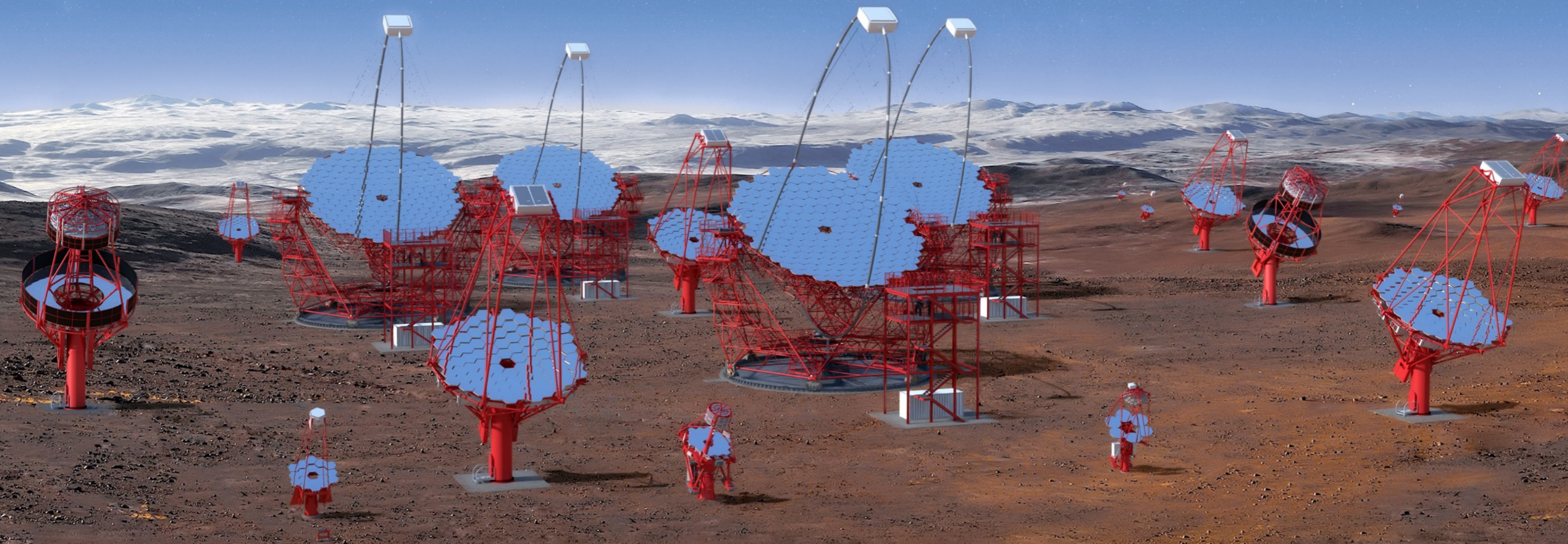


Image Credit: G. Pérez, IAC, SMM

Cherenkov Telescope Array (CTA)



cherenkov
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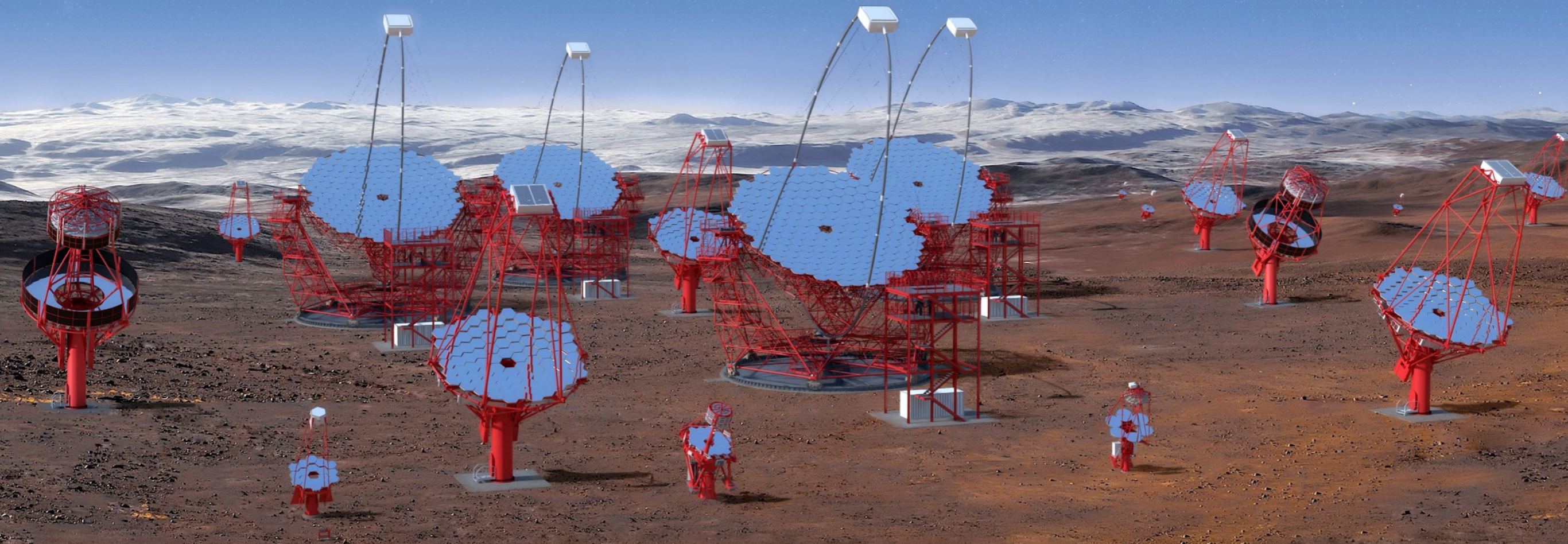


Image Credit: G. Pérez, IAC, SMM

Cherenkov Telescope Array (CTA)



cherenkov
telescope
array

Large-Sized Telescope (LST)

Dia. : 23 m

Energy : 20–150 GeV

N_{Tel} : 4 @ North, 4 @ South

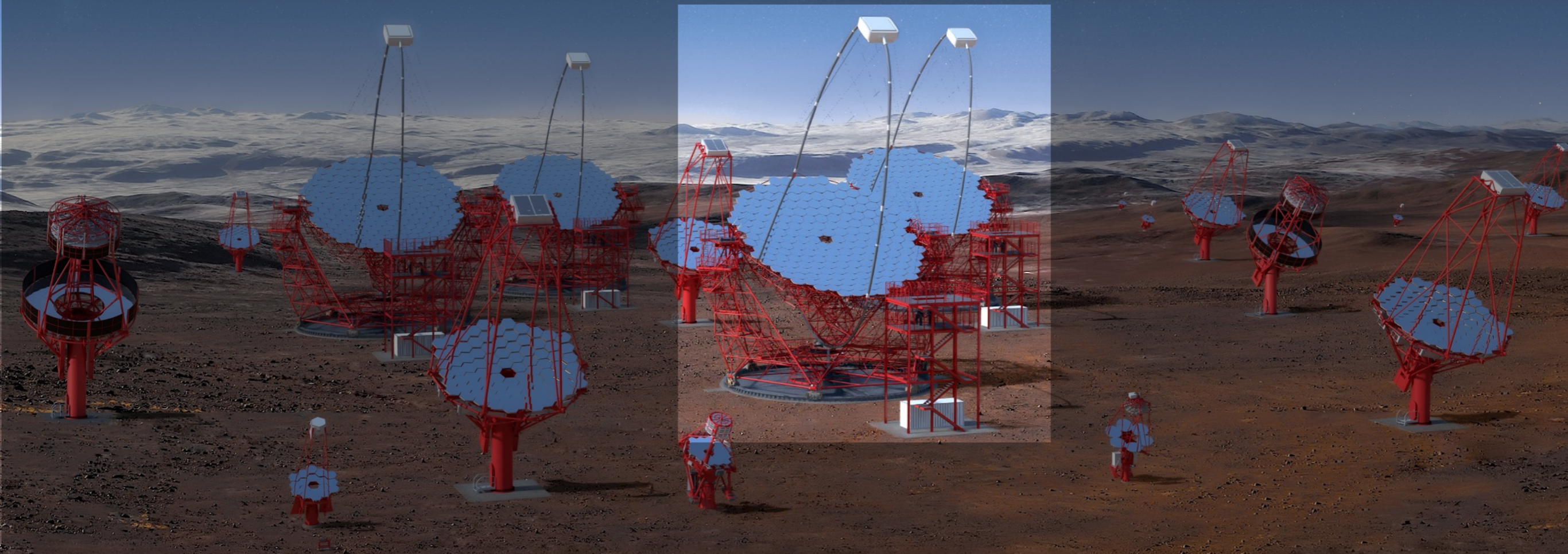


Image Credit: G. Pérez, IAC, SMM

Cherenkov Telescope Array (CTA)



cherenkov
telescope
array

Medium-Sized Telescope (MST)

Dia. : 12 m

Energy : 150 GeV–5 TeV

N_{Tel} : 15 @ North, 25 @ South

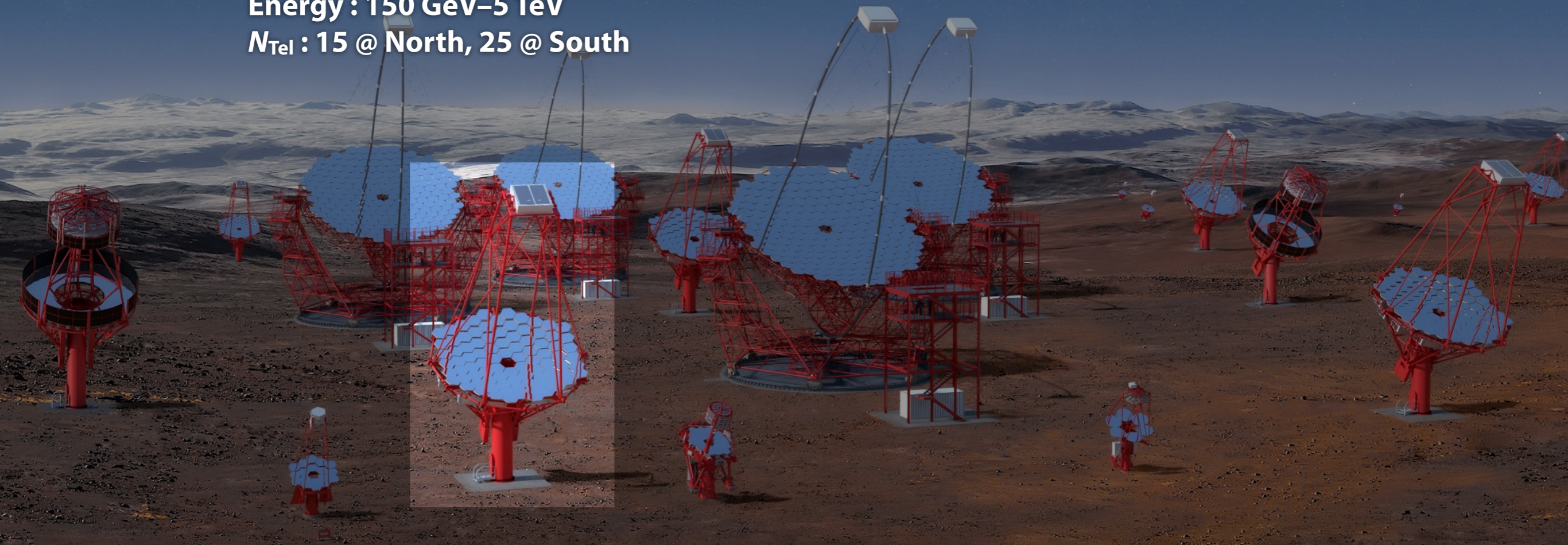
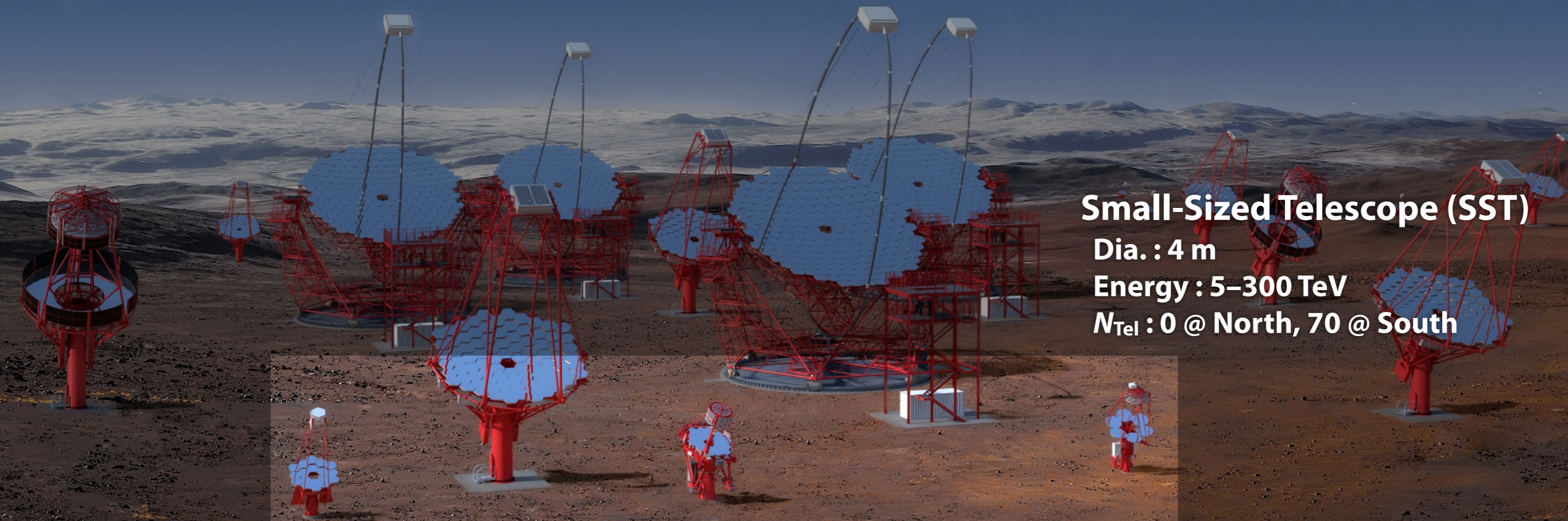


Image Credit: G. Pérez, IAC, SMM

Cherenkov Telescope Array (CTA)



cherenkov
telescope
array



Small-Sized Telescope (SST)

Dia. : 4 m

Energy : 5–300 TeV

N_{Tel} : 0 @ North, 70 @ South

Image Credit: G. Pérez, IAC, SMM

Cherenkov Telescope Array (CTA)



cherenkov
telescope
array

Schwarzschild–Couders Telescope (SCT)

Dia. : 10 m

Energy : 150 GeV–5 TeV

N_{Tel} : 15 @ North, 25 @ South (incl. MSTs)

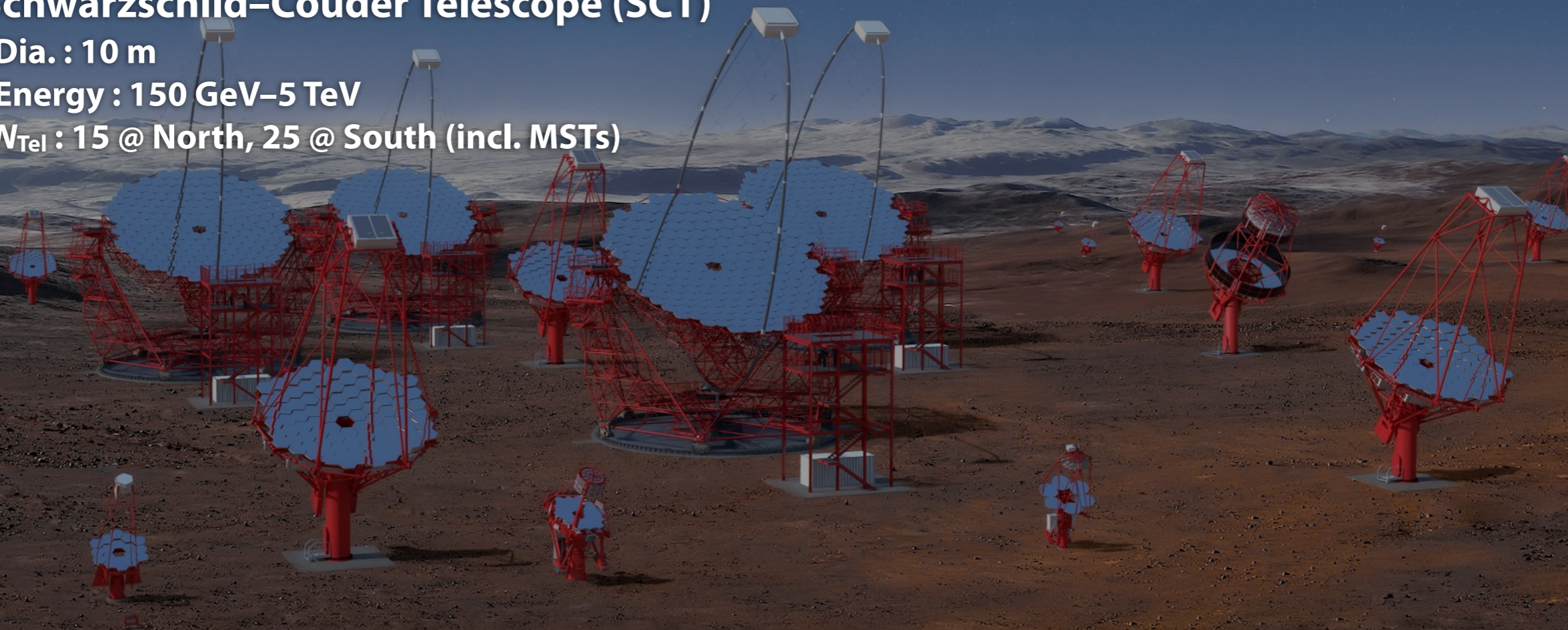
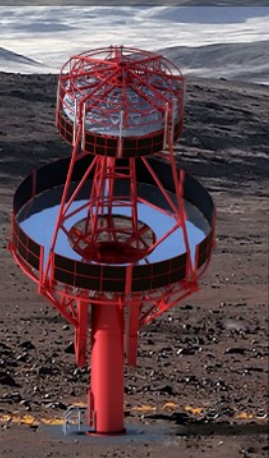
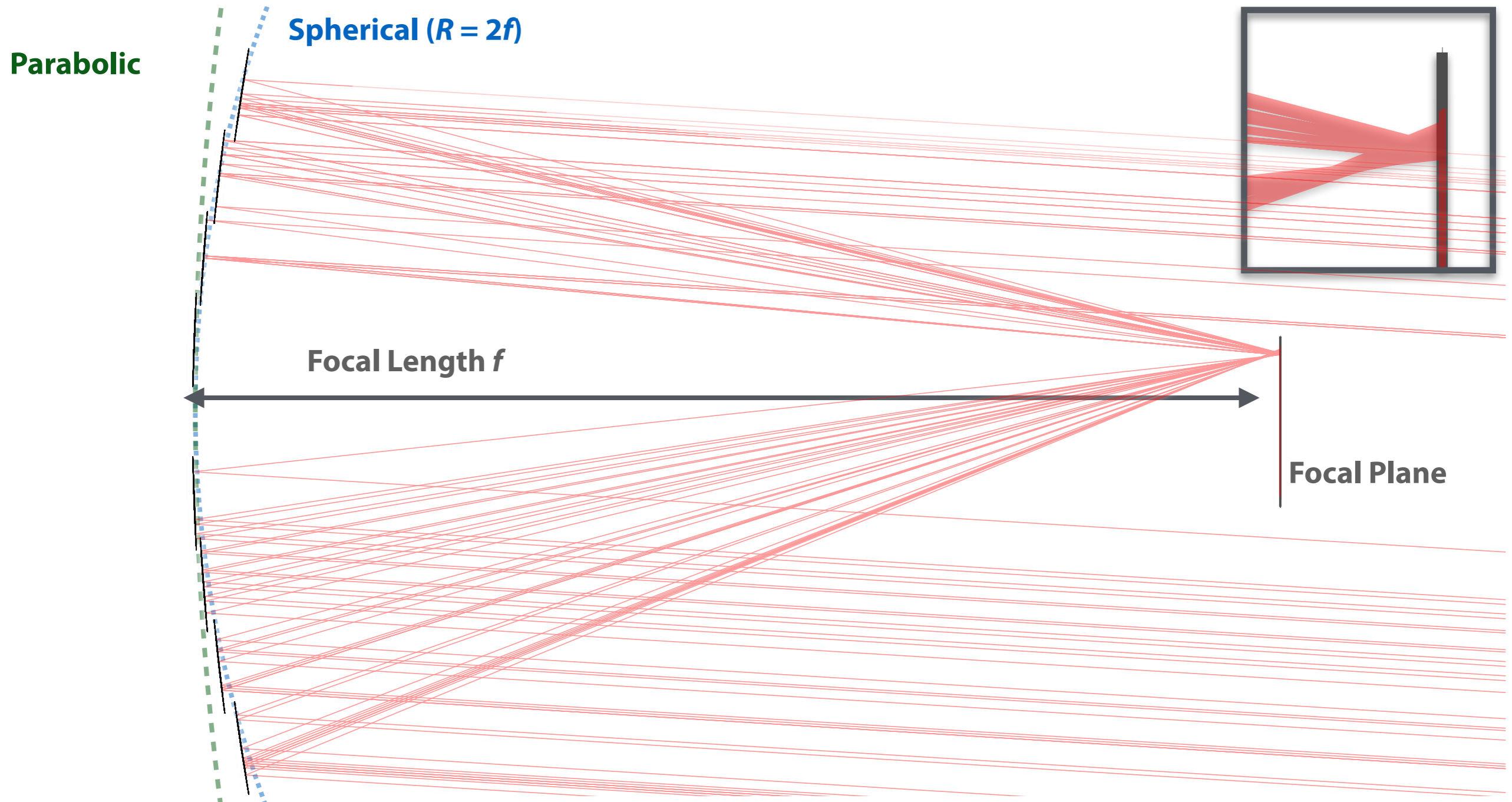


Image Credit: G. Pérez, IAC, SMM

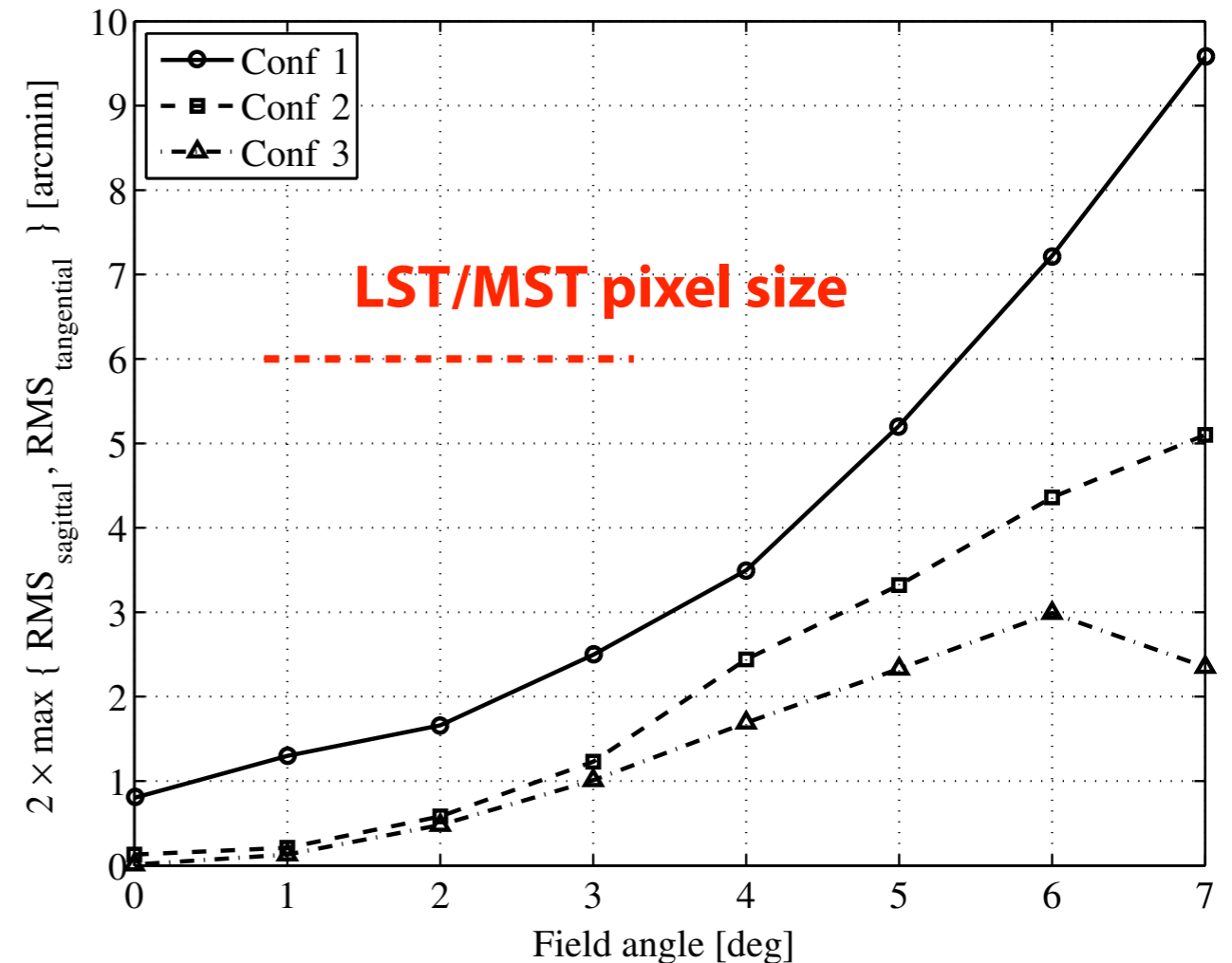
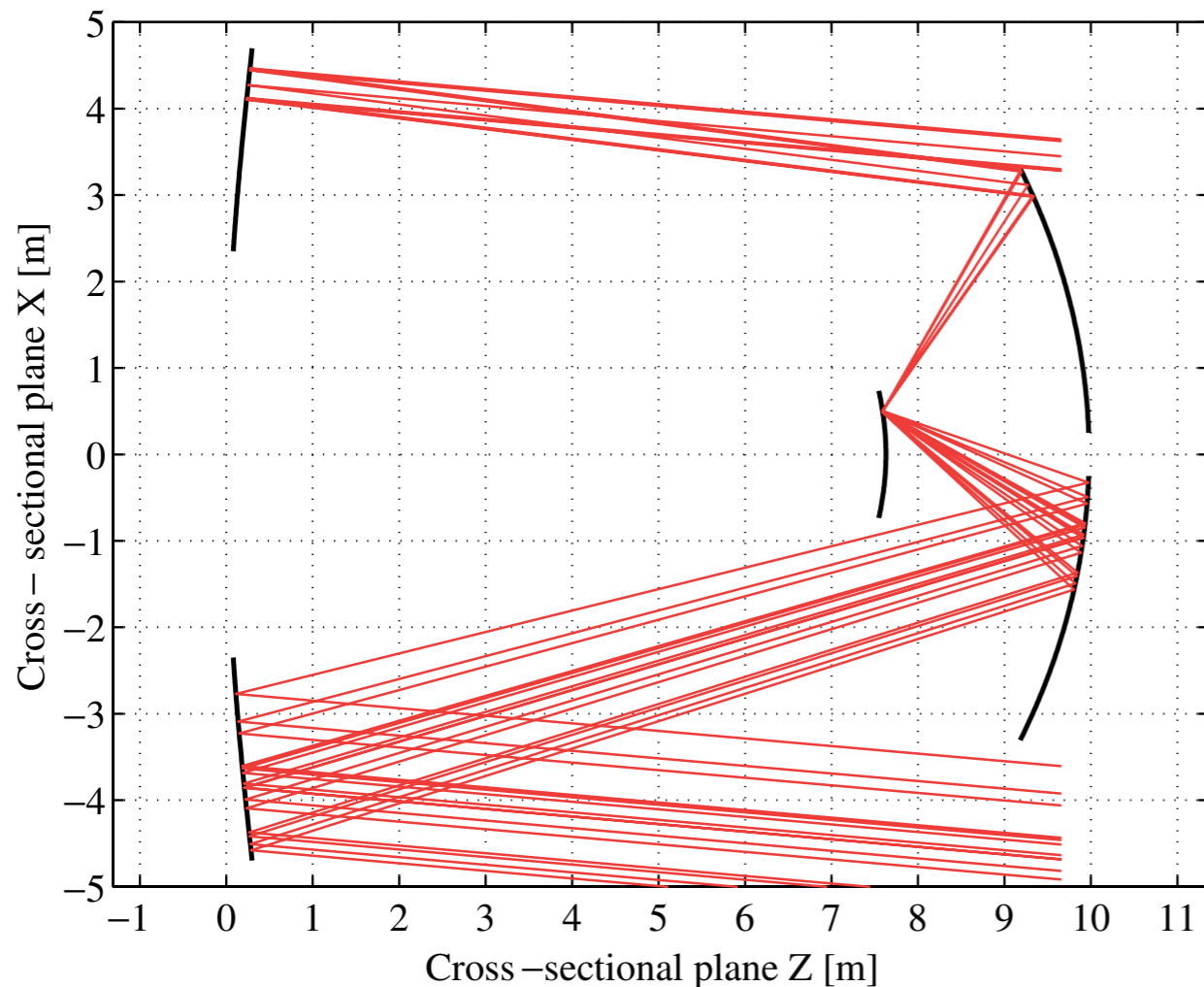
Davies–Cotton Configuration



- Another standard optics design in Cherenkov telescopes (Parabola for LSTs)
- Initially proposed for solar power plants
- Wide field-of-view (FOV), but thus low angular resolution

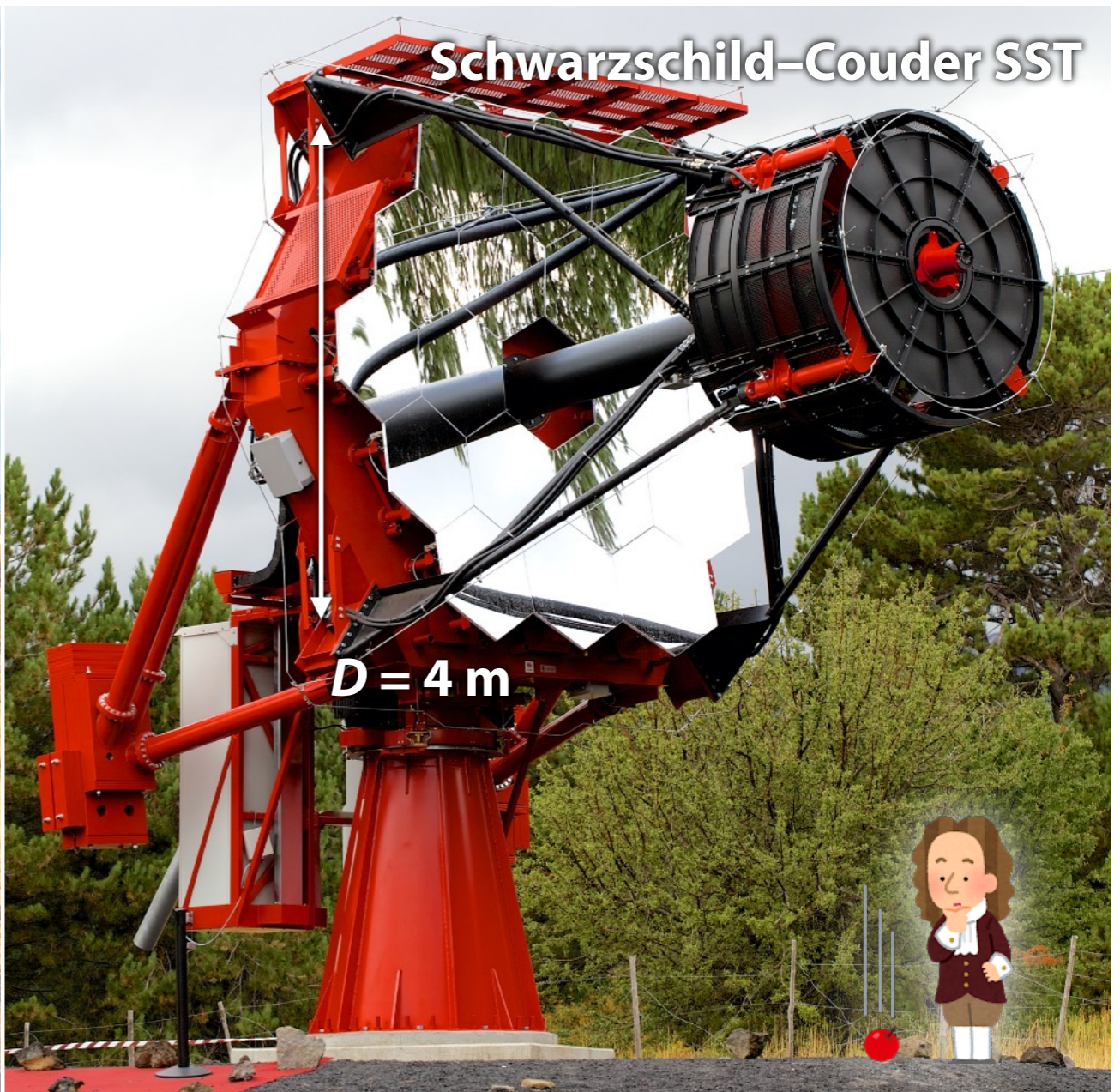
Schwarzschild–Couder Configuration

Vassiliev+ (2007)



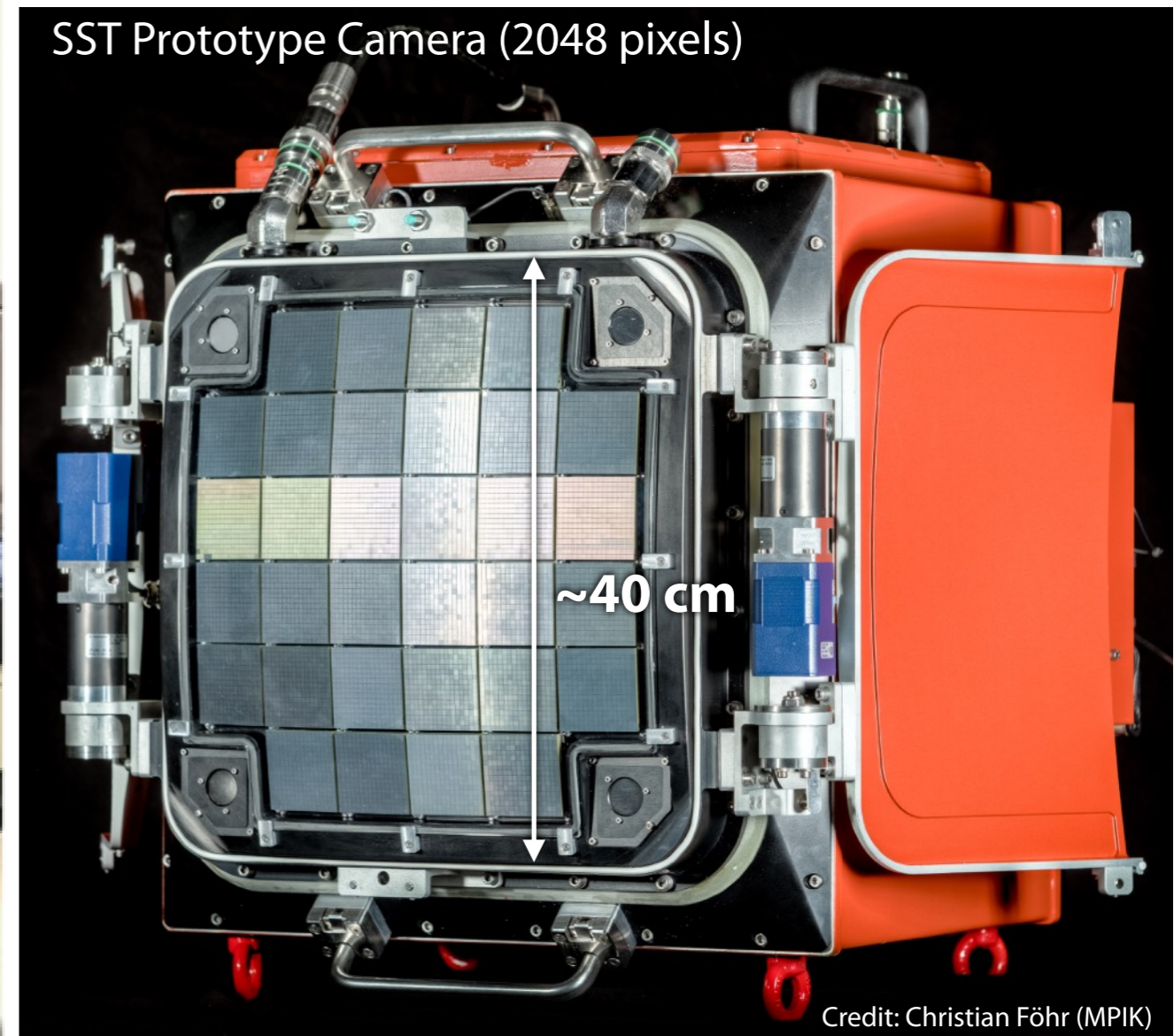
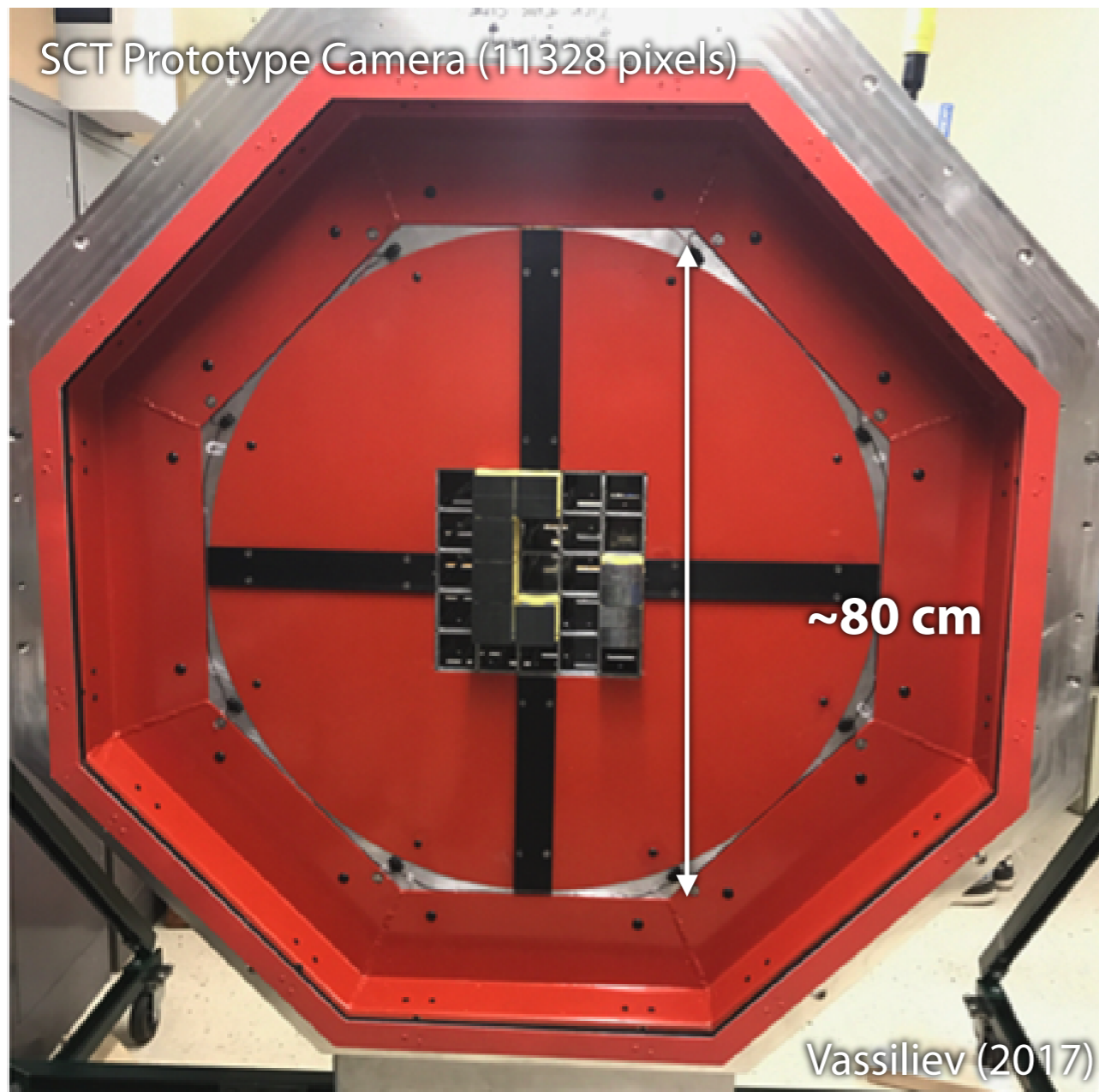
- **Aspherical** primary and **secondary** mirrors to achieve wide FOV and better resolution at the same time
- Wider FOV brings fast survey and wider effective area for higher-energy photons
- Finer shower-image resolution (→ higher sensitivity) and compact camera (→ less expensive) are expected
- Initially proposed by the CTA US group for MSTs

Schwarzschild–Couder Proposals for CTA

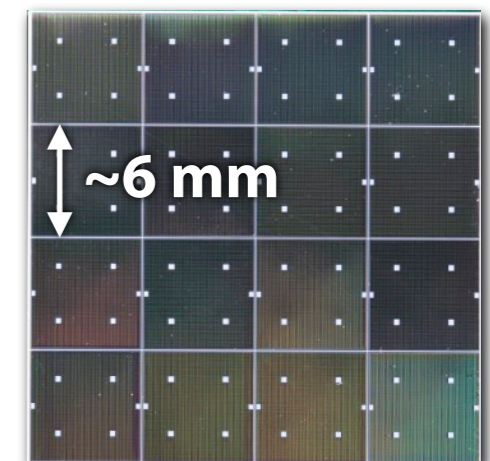


- Schwarzschild–Couder MST (SCT) with 10-meter diameter for MST extension
 - ▶ US, Italy, Germany, Mexico, and Nagoya
 - ▶ Japanese group contributions in SiPM, electronics, MC, and software
- 4-meter SC SST × camera design chosen to be the final SST design from three proposals in June 2019
 - ▶ Similar Japanese contributions but more active in SST

Need Compact Cameras with SiPMs



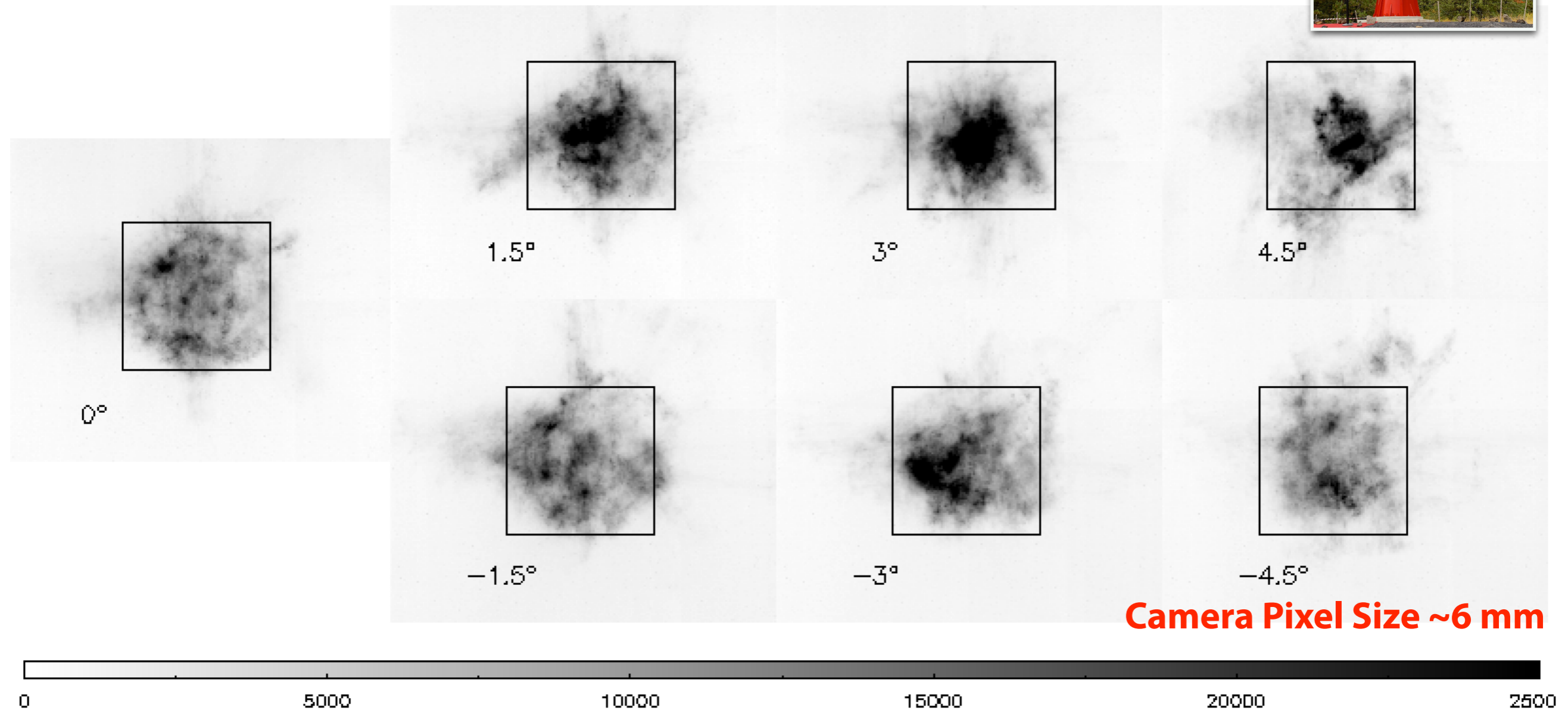
- The concave secondary mirrors make the plate scales ($\equiv 1/f$) large and enable us to build compact cameras
- Silicone photomultipliers (SiPMs) are used instead of conventional photomultiplier tubes (PMTs)



SST Optical System

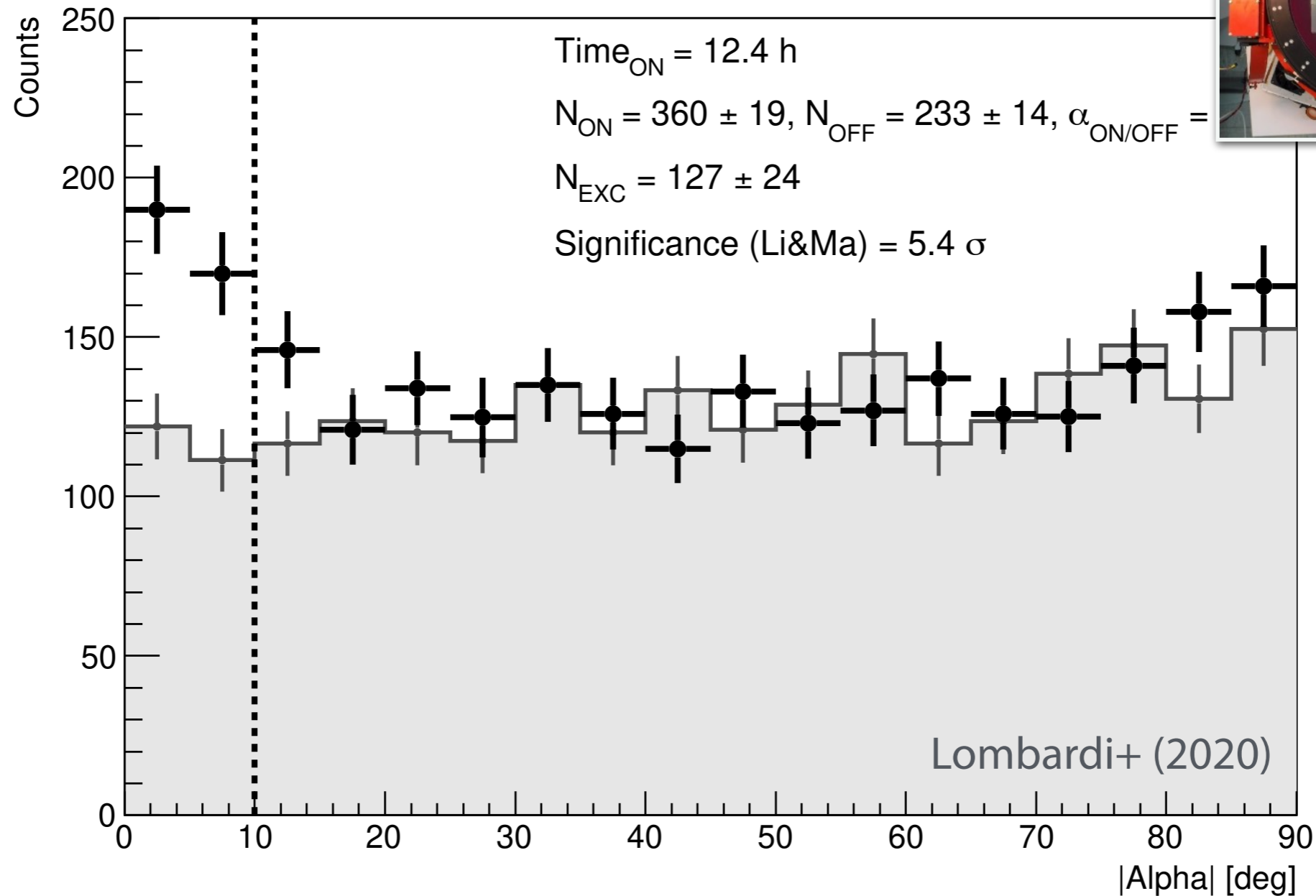
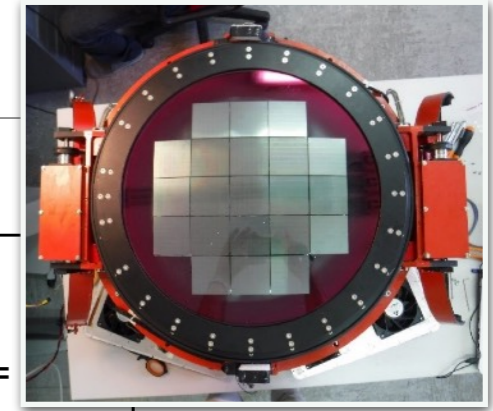


Giro+ (2017)



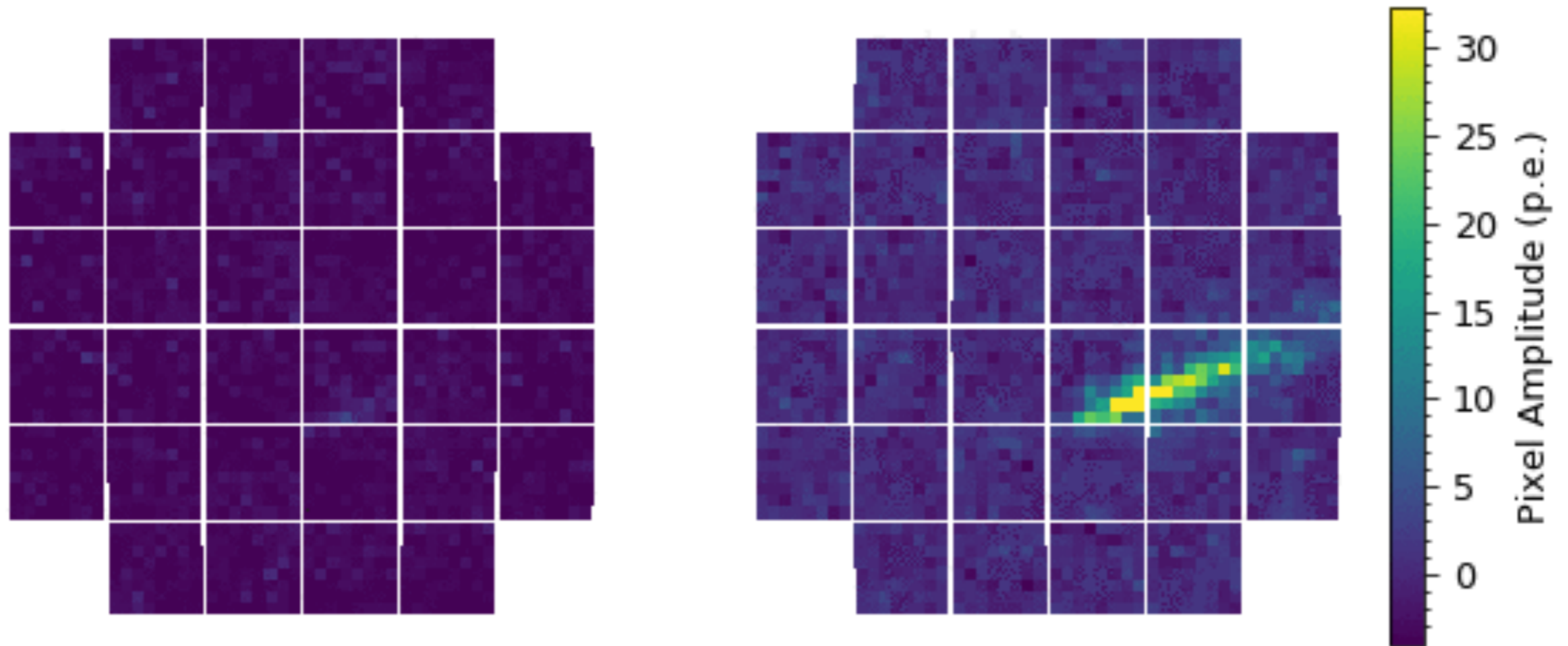
- Achieved good enough optical resolution matching the SST pixel size
- First realization of the Schwarzschild–Couder configuration ever

SST Crab Observations: Detection at 5.4σ



- “ASTRI” camera (lead by Italian groups) successfully detected gamma-ray signal from the Crab Nebula
- Combination of the Schwarzschild–Couder and a SiPM camera verified

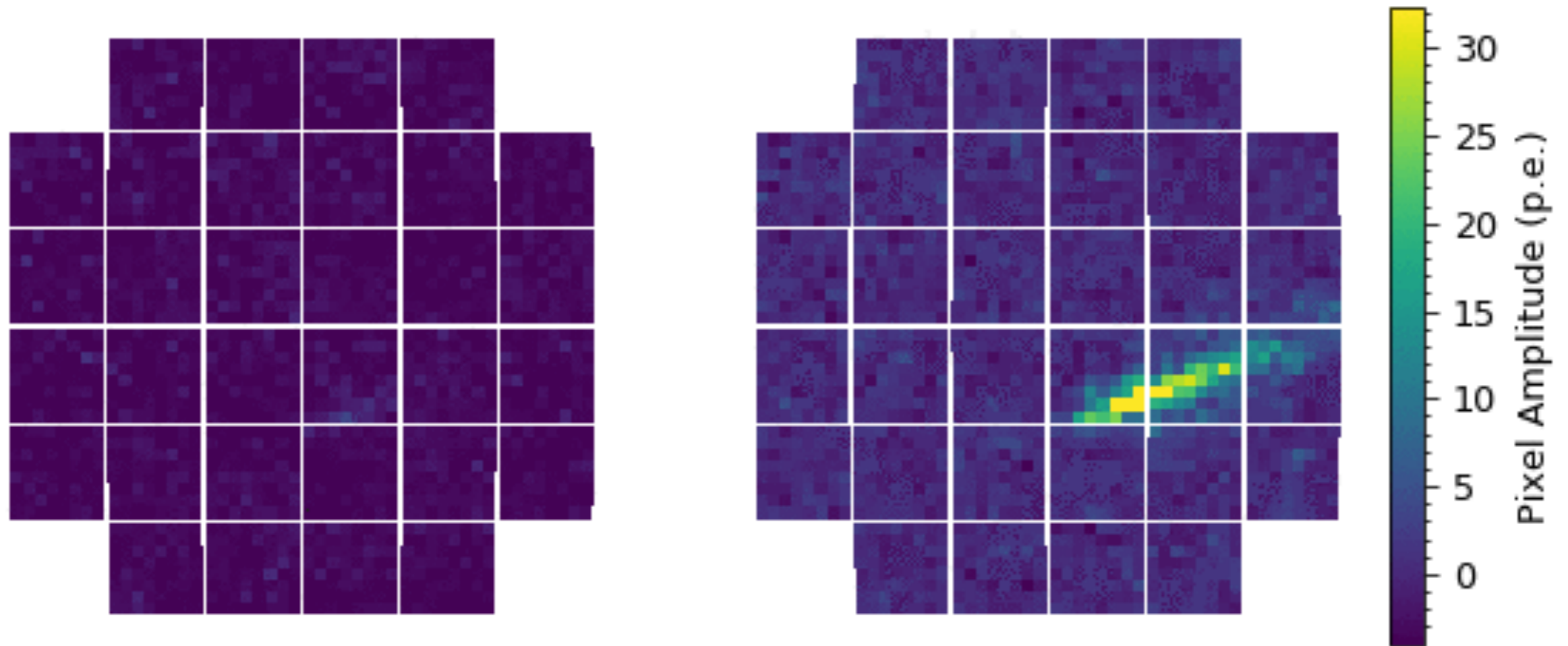
Another (Our) SST Camera Test Observations



<https://www.cta-observatory.org/chec-achieves-first-light-on-astri/>

- “Our” camera also succeeded in air-shower observations on the same prototype telescope (replaceable with the Italian camera)
- Additional test observations canceled due to mirror re-coating and COVID-19

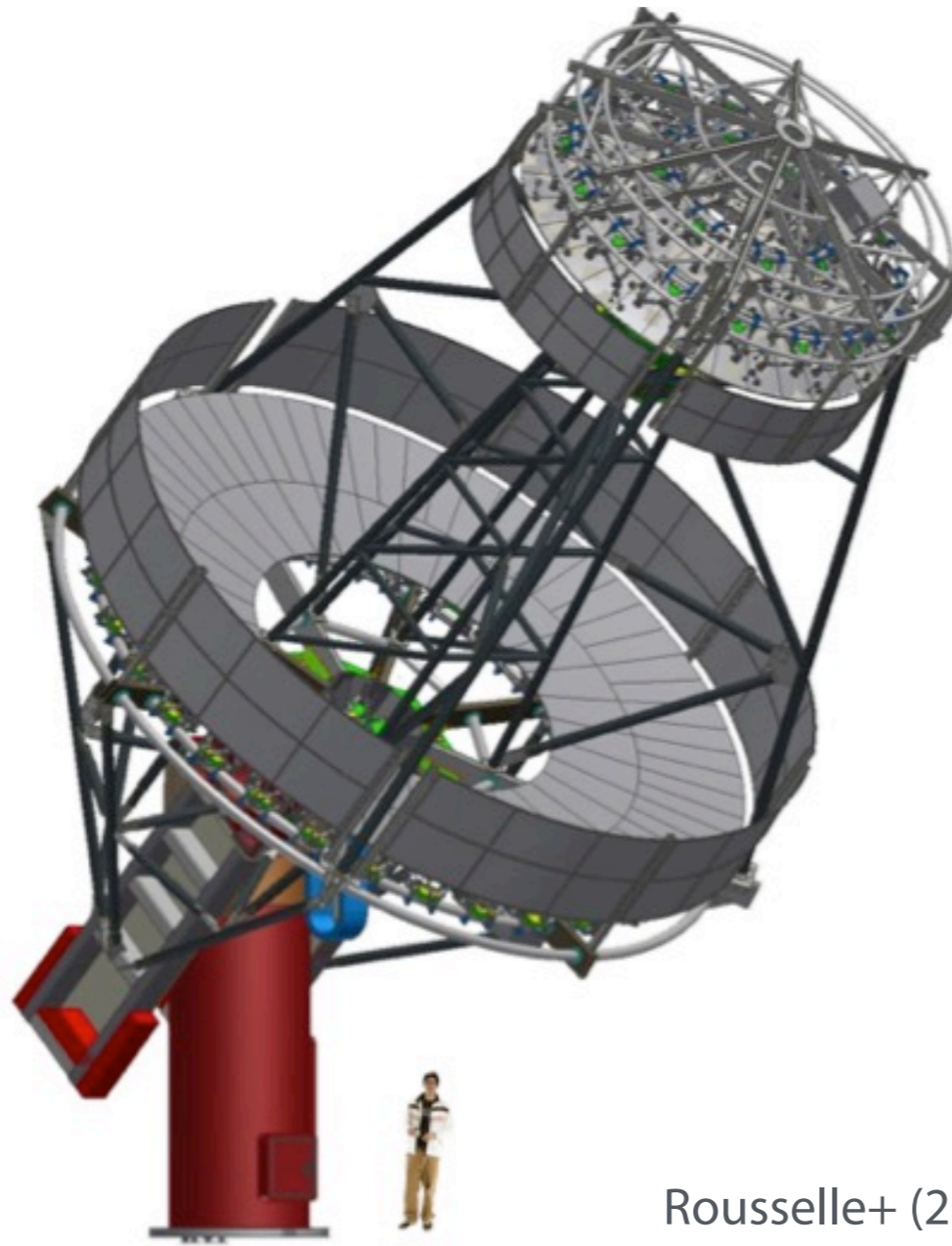
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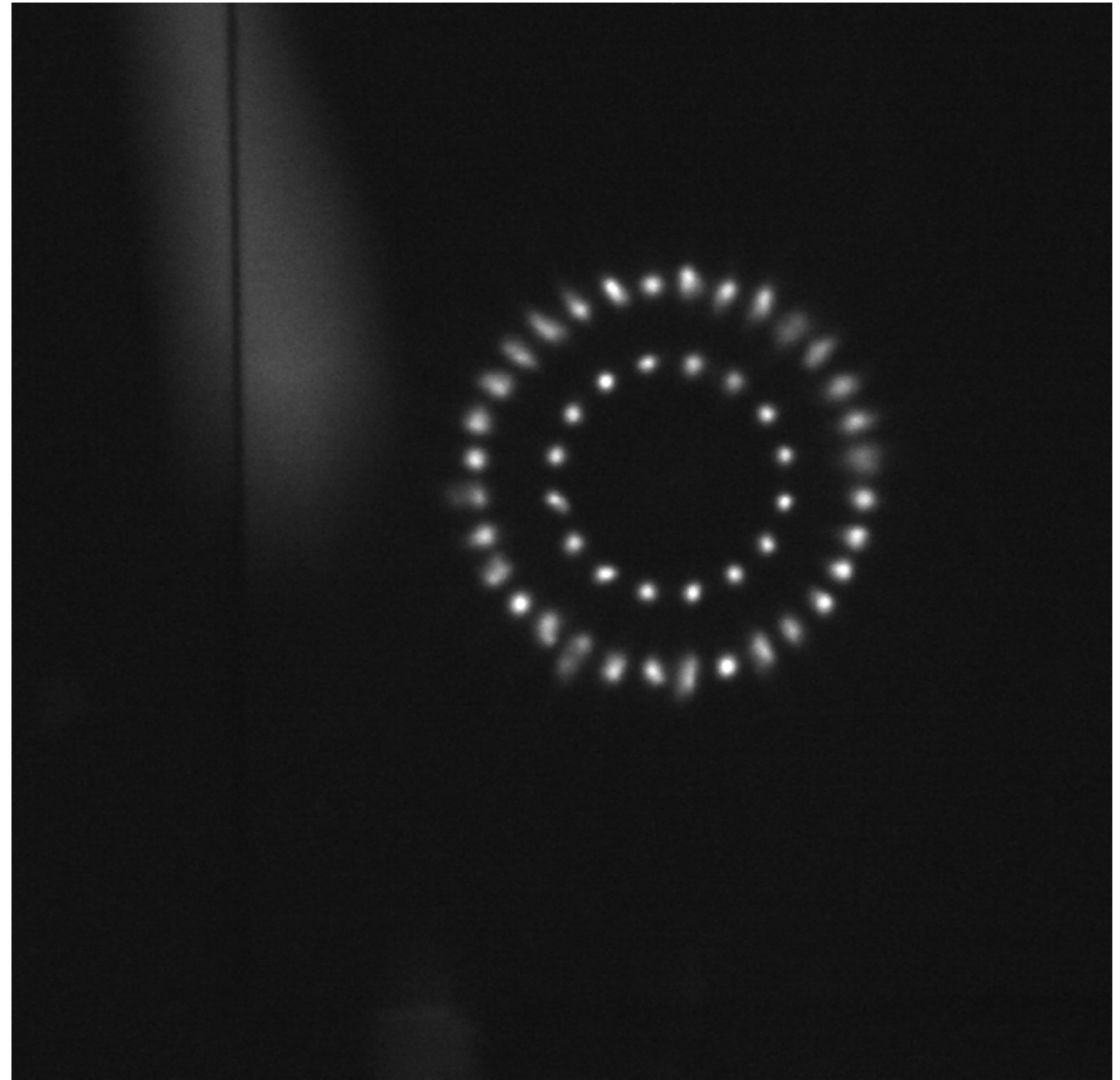
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Schwarzschild–Couder MST Optical System

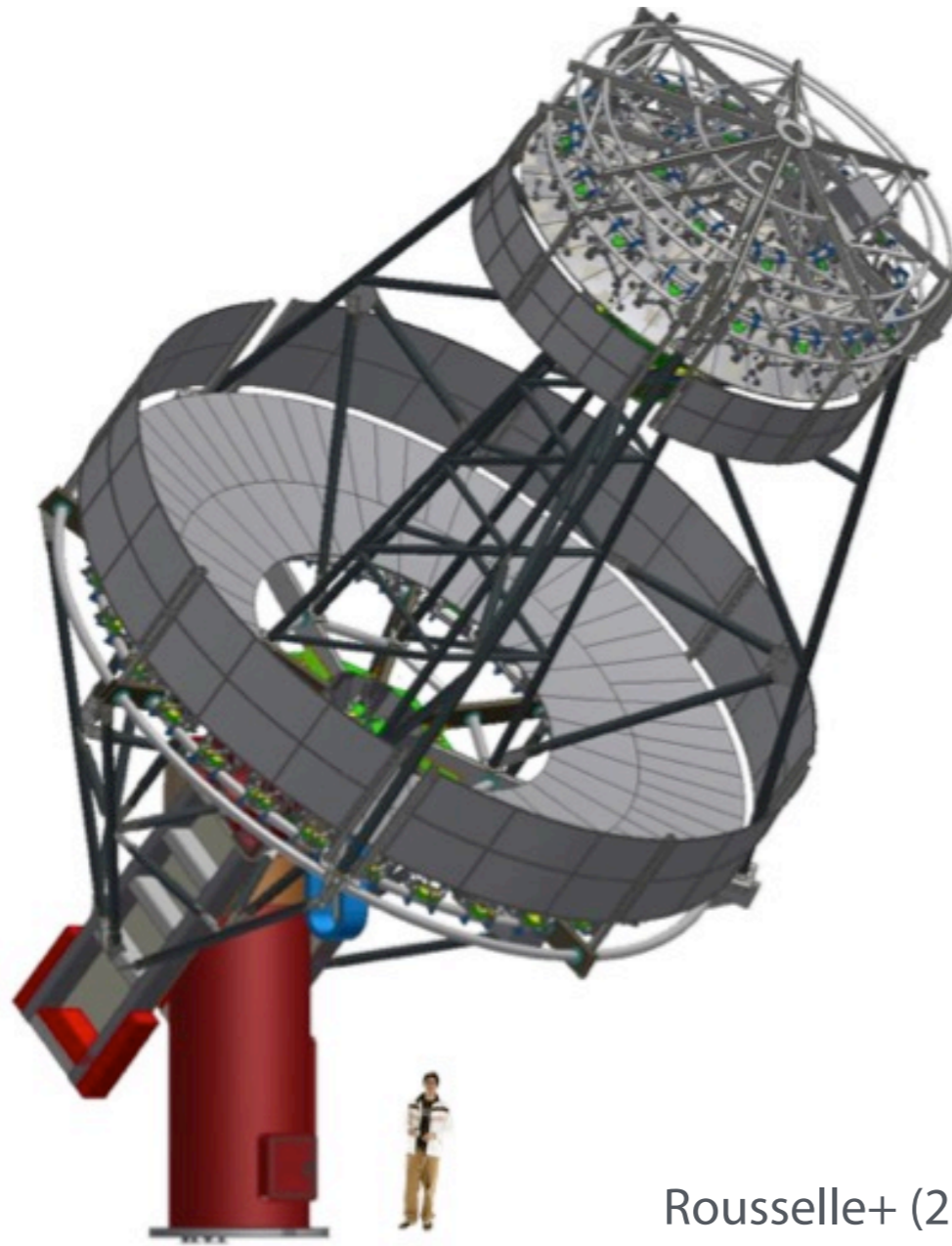


Rousselle+ (2015)

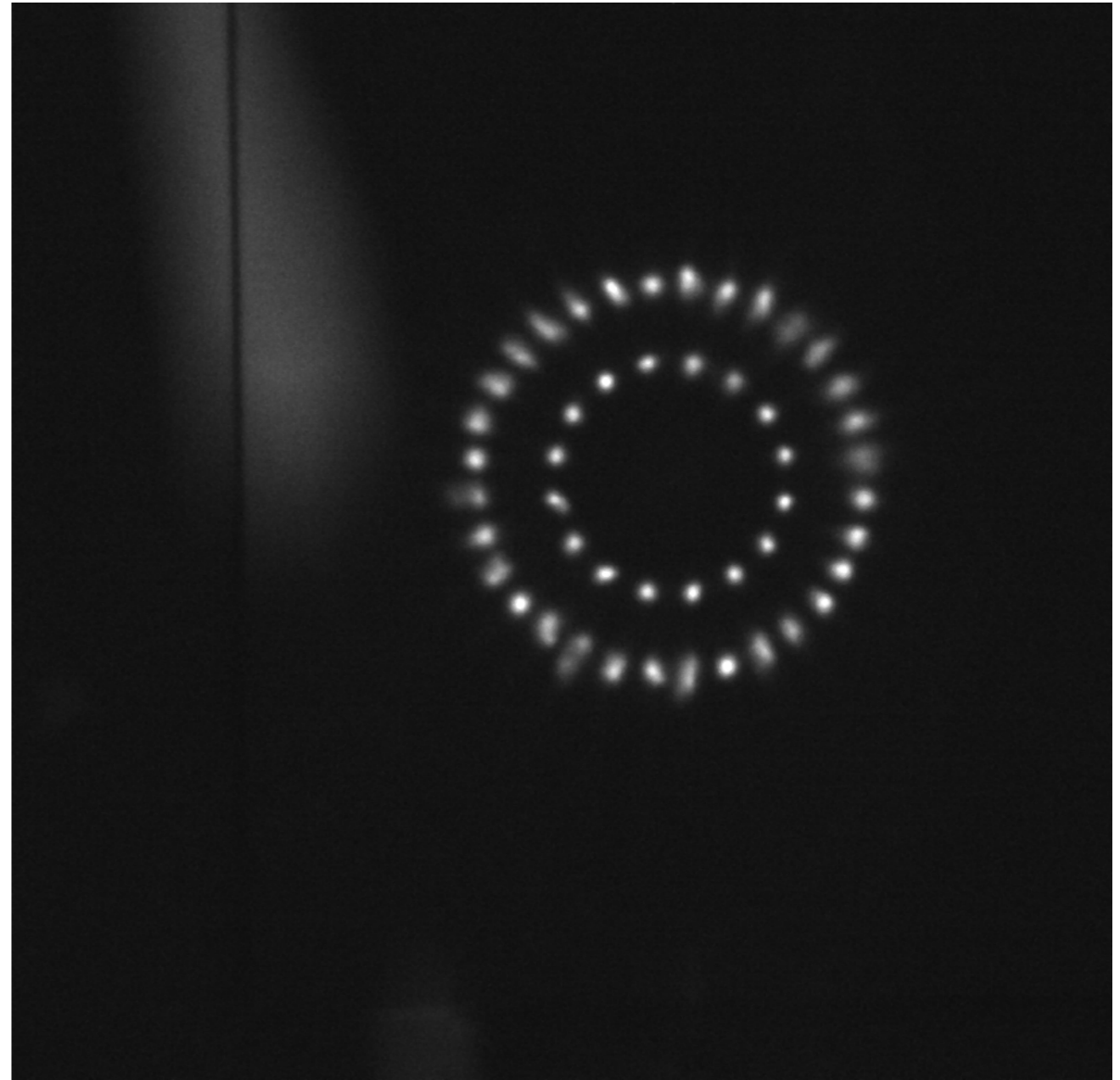


- 2.5 times larger than the SST optical system
- More number of segmented mirrors, thus more complex
- Optical alignment was successfully finished

Schwarzschild–Couder MST Optical System



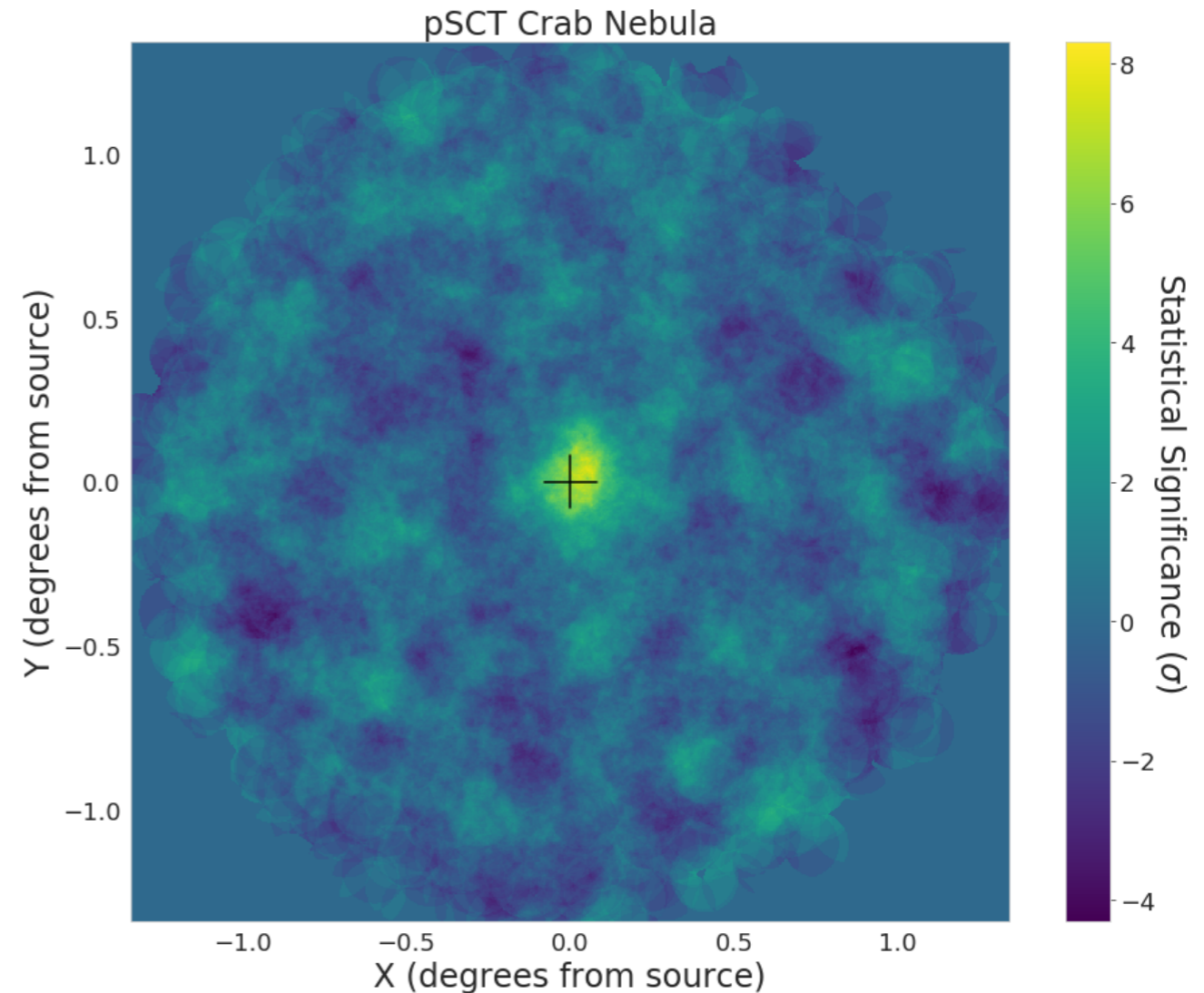
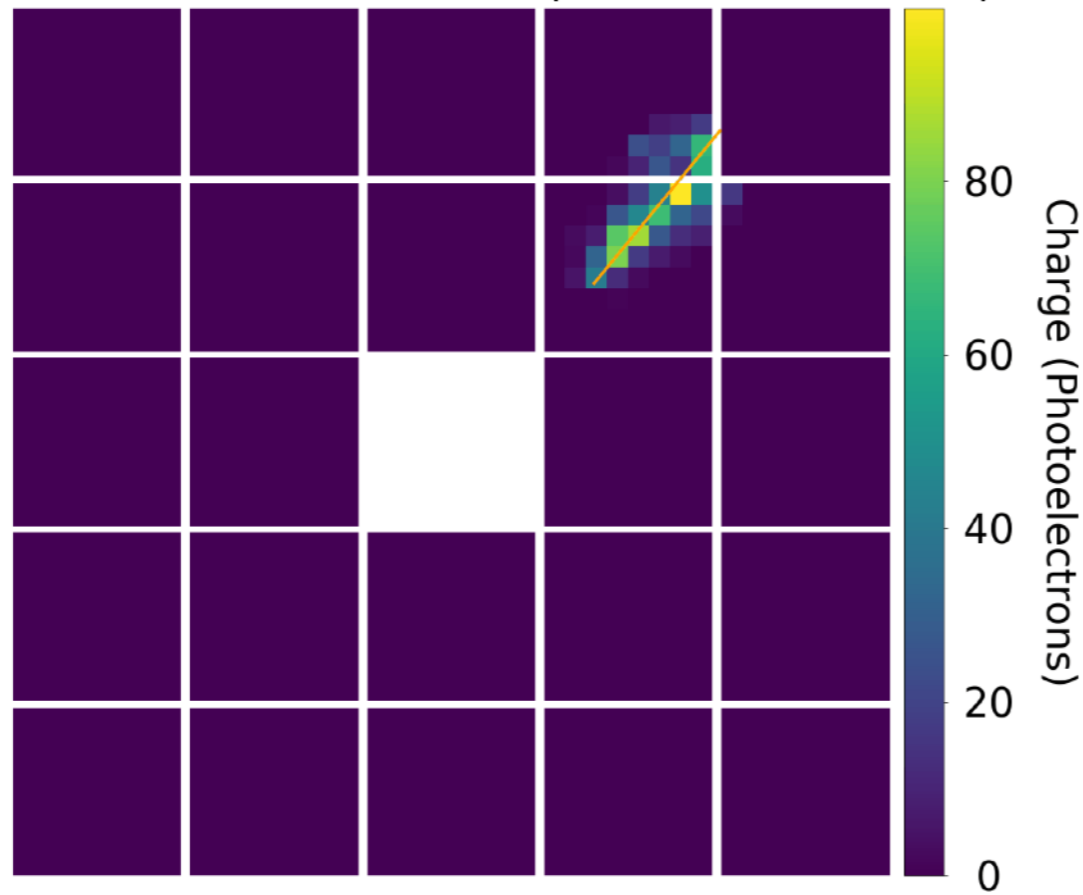
Rousselle+ (2015)



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Crab Observations by Prototype Schwarzschild–Couder

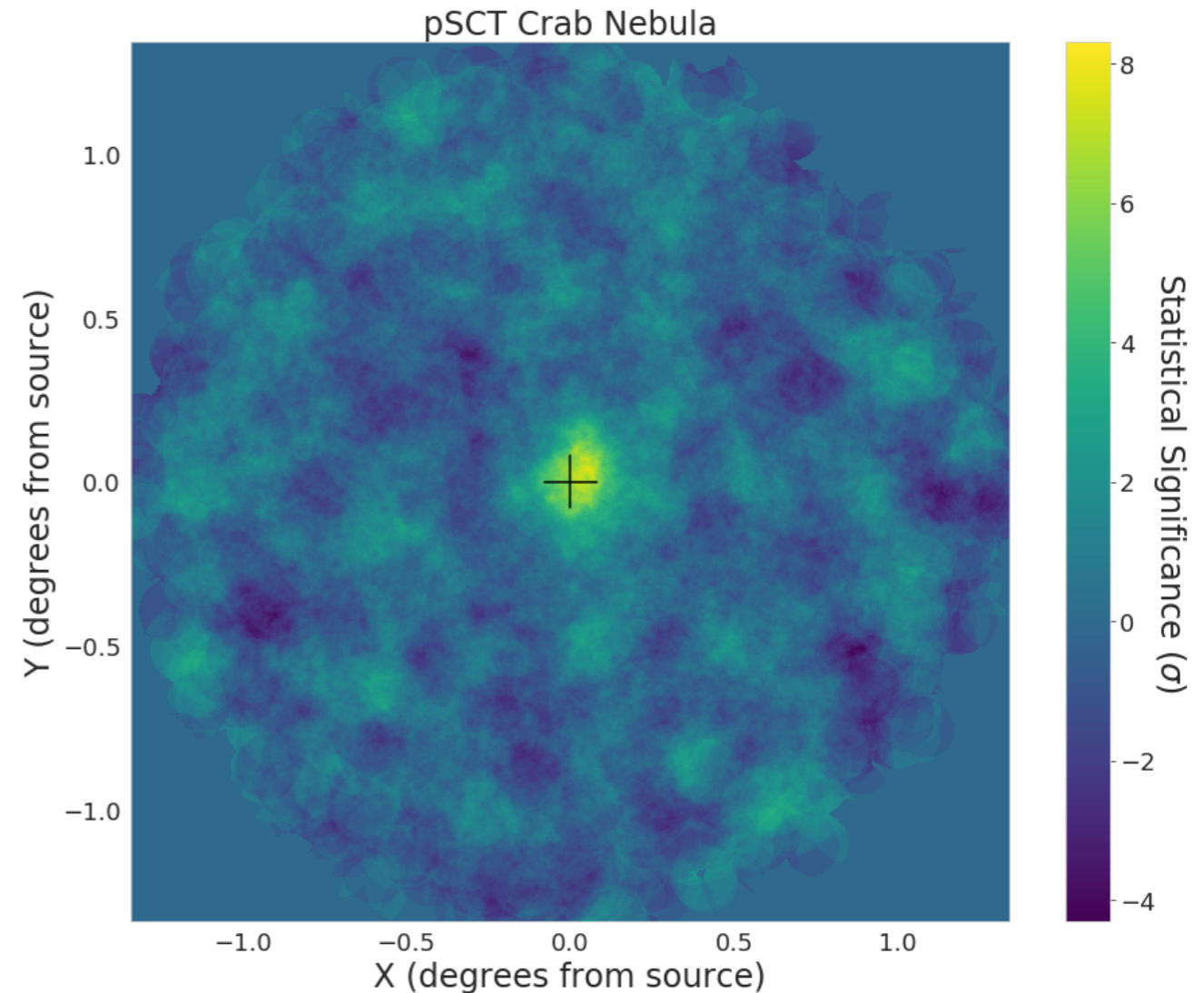
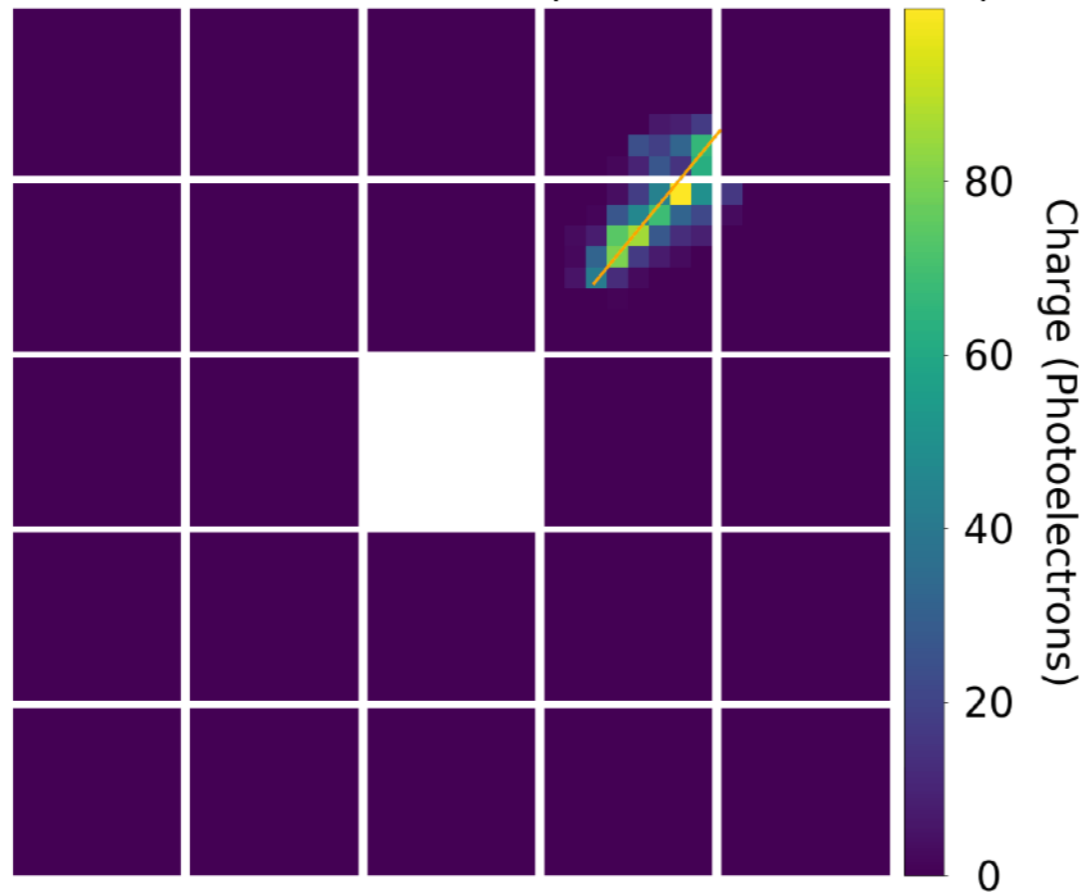
Prototype Schwarzschild-Couder Telescope Gamma Rays
Run 328629 Event 085862 (2020-01-28 04:22:10)



- Crab Nebular detection at 8.3σ , while the prototype FOV is still limited

Crab Observations by Prototype Schwarzschild–Couder

Prototype Schwarzschild-Couder Telescope Gamma Rays
Run 328629 Event 085862 (2020-01-28 04:22:10)



- Crab Nebular detection at 8.3σ , while the prototype FOV is still limited

Summary and Plans

- Prototype Schwarzschild–Couder (SC) telescopes for the CTA Medium- and Small-sized Telescopes constructed
- First realization of the SC configuration
- Both succeeded in air-shower Cherenkov observations and Crab Nebula detection with prototype SiPM cameras

- Prototype SC-MST will be upgraded with full FOV coverage with more SiPM tiles and new electronics
- SST is going to finalize the optics and camera designs
- New SST prototype(s) will be built and tested before the pre-production phase