

Schwarzschild–Couder 光学系を用いた CTA 小・中口径望遠鏡の開発

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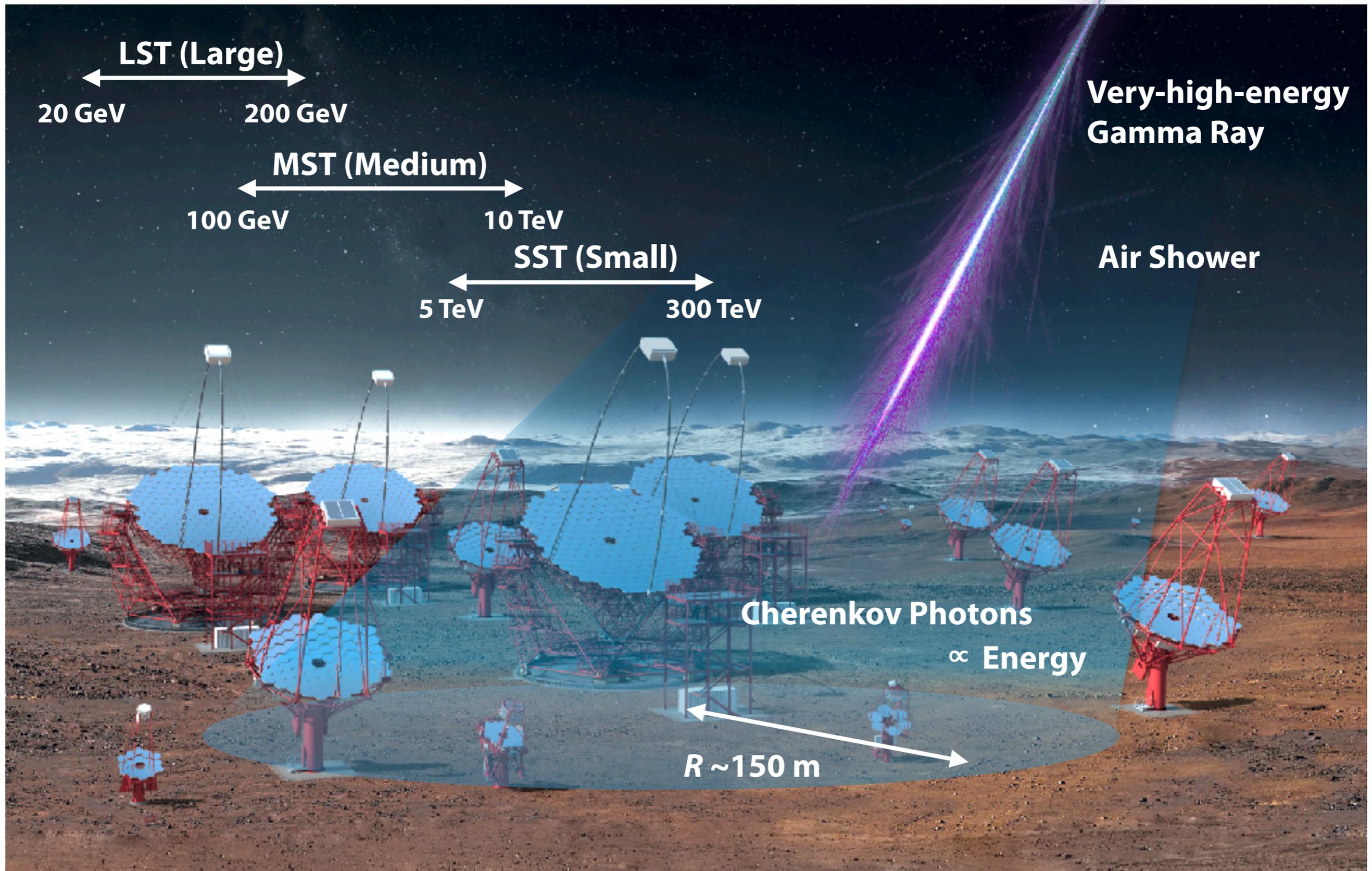
² 茨城大学理学部

日本天文学会 2017 年 秋季年会

Cherenkov Telescope Array (CTA)



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Cherenkov Telescope Array (CTA)

Small-Sized Telescope (SST)

70 SSTs @ South

$D = 4 \text{ m}$

FOV $\sim 9^\circ$

$E = 5 \text{ TeV} - 300 \text{ TeV}$

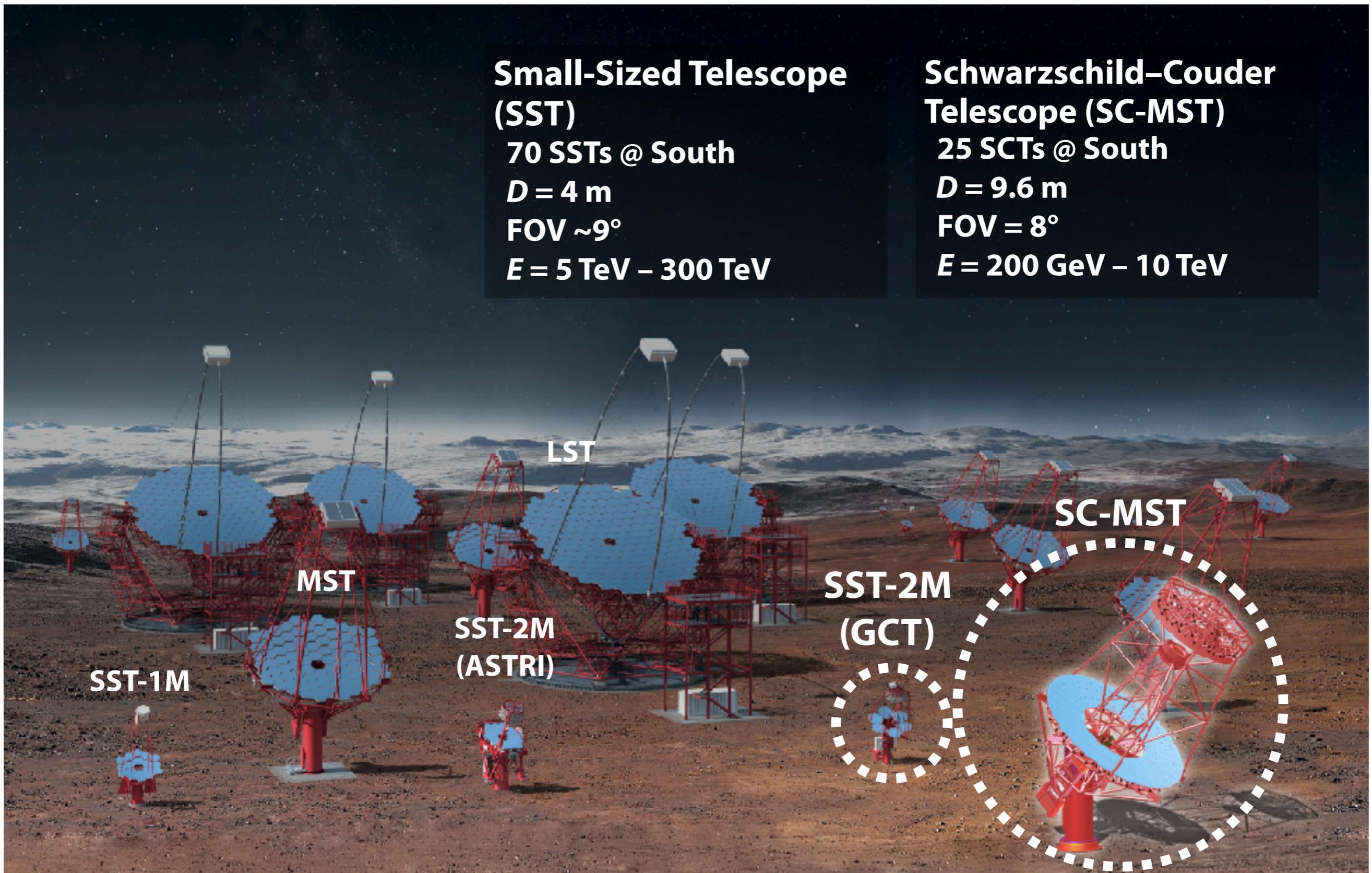
Schwarzschild-Couder Telescope (SC-MST)

25 SCTs @ South

$D = 9.6 \text{ m}$

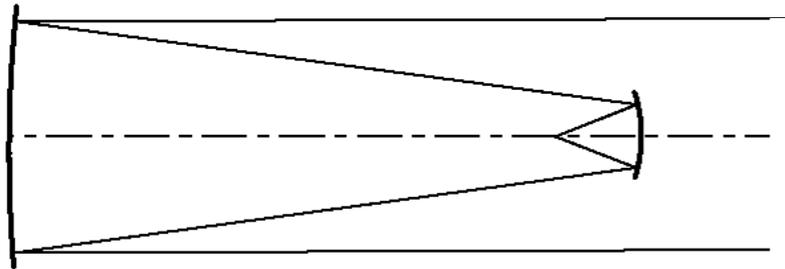
FOV = 8°

$E = 200 \text{ GeV} - 10 \text{ TeV}$

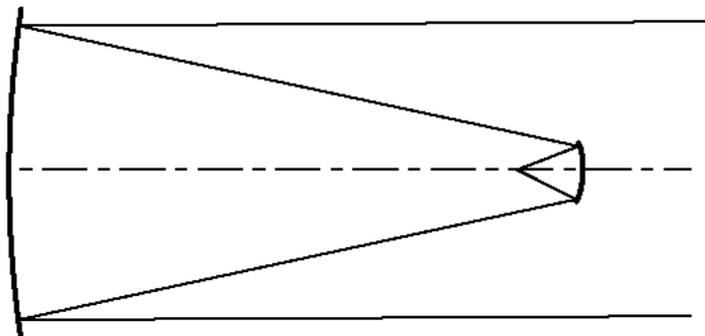


The Schwarzschild–Couder (SC) Design

“Handbook of Optics”

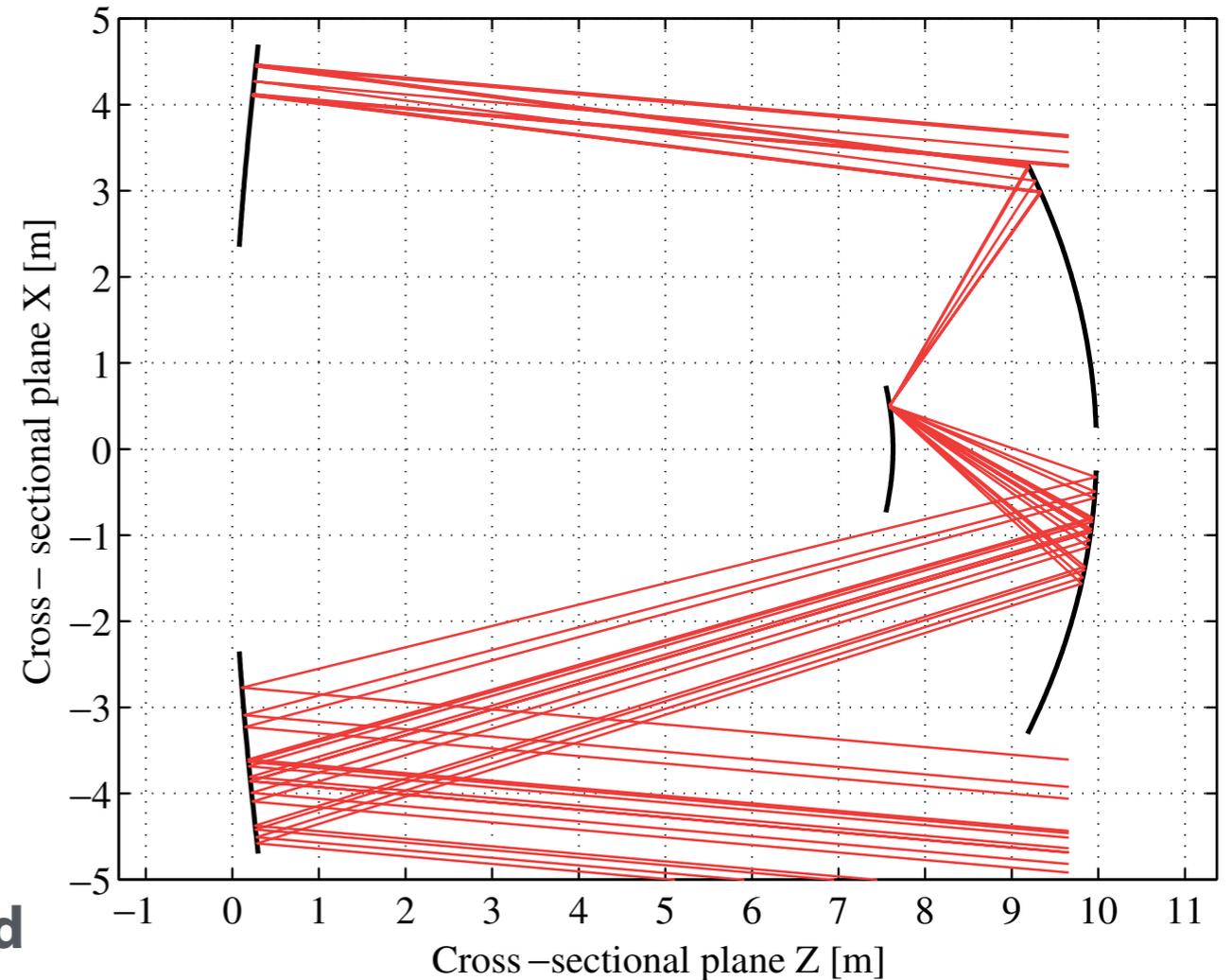


Couder (1926)



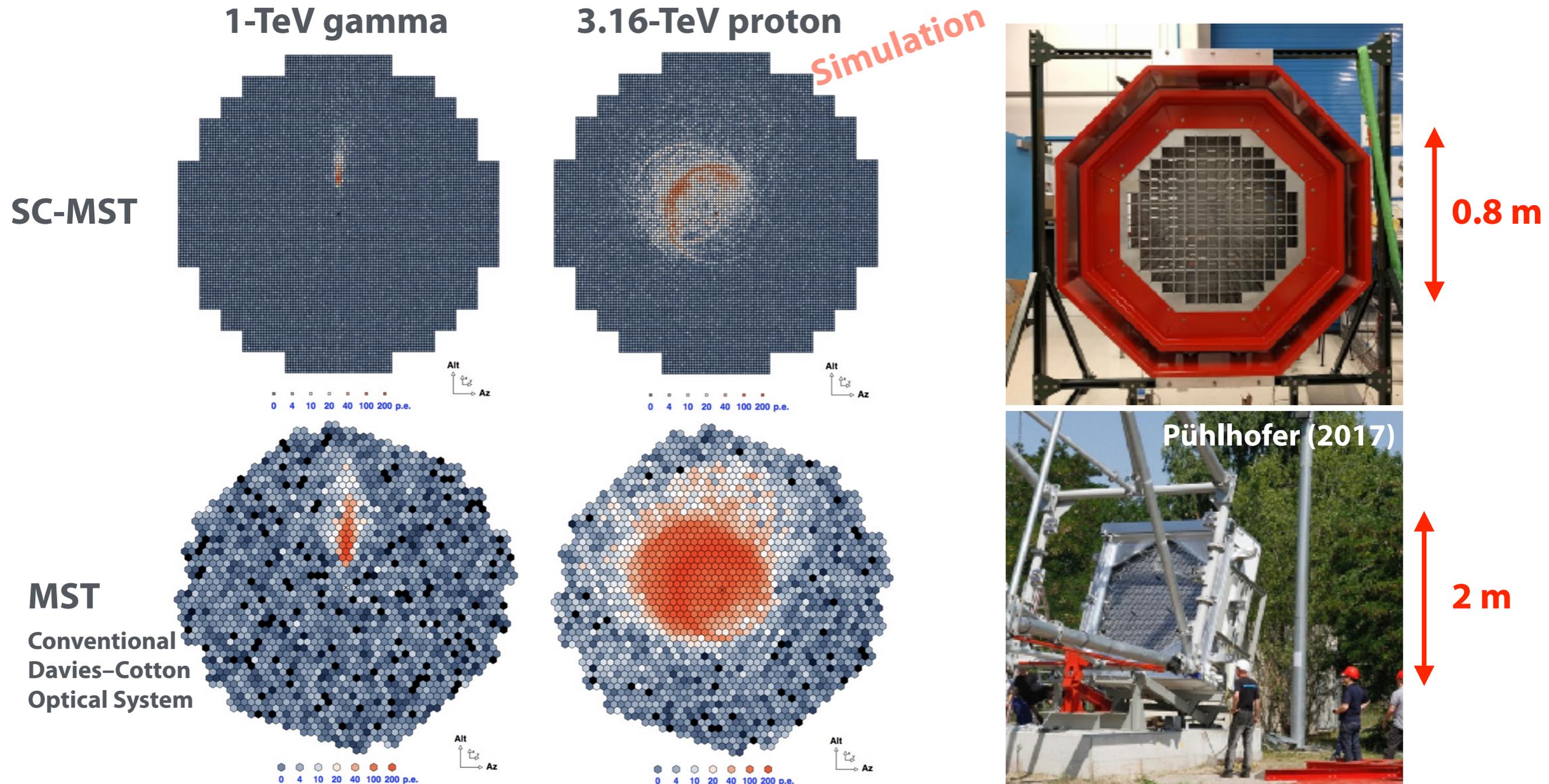
Schwarzschild
(1905)

Vassiliev *et al.* (2007)



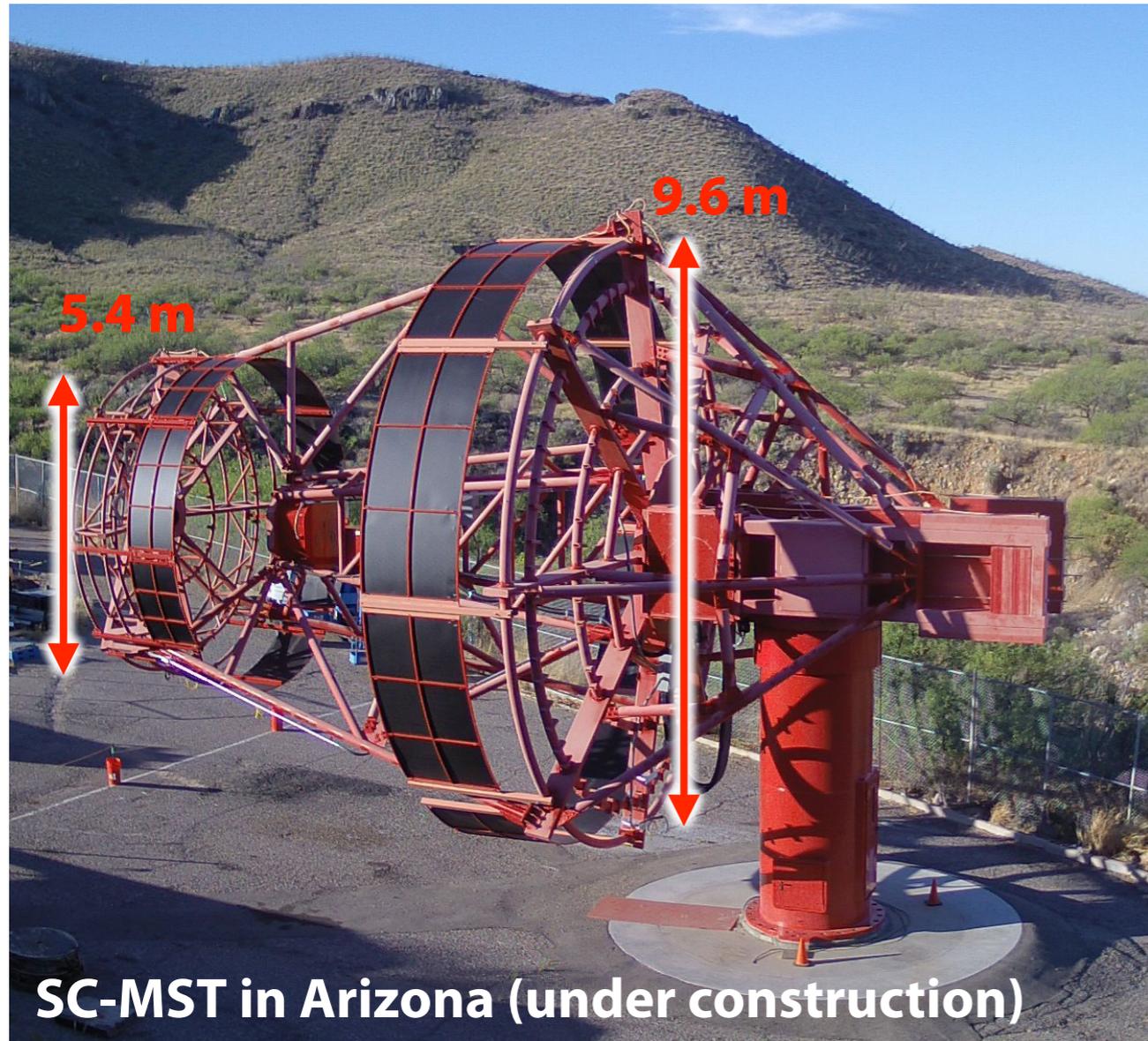
- Wide FOV aplanatic design with primary and secondary mirrors, invented by Schwarzschild (1905) and Couder (1926)
- Proposed for ground-based gamma-ray telescopes in 2007
- Will achieve **wider FOV** ($\sim 8^\circ$) and **higher resolution** ($< \sim 0.04^\circ$) with a compact camera

Improved Optical Resolution



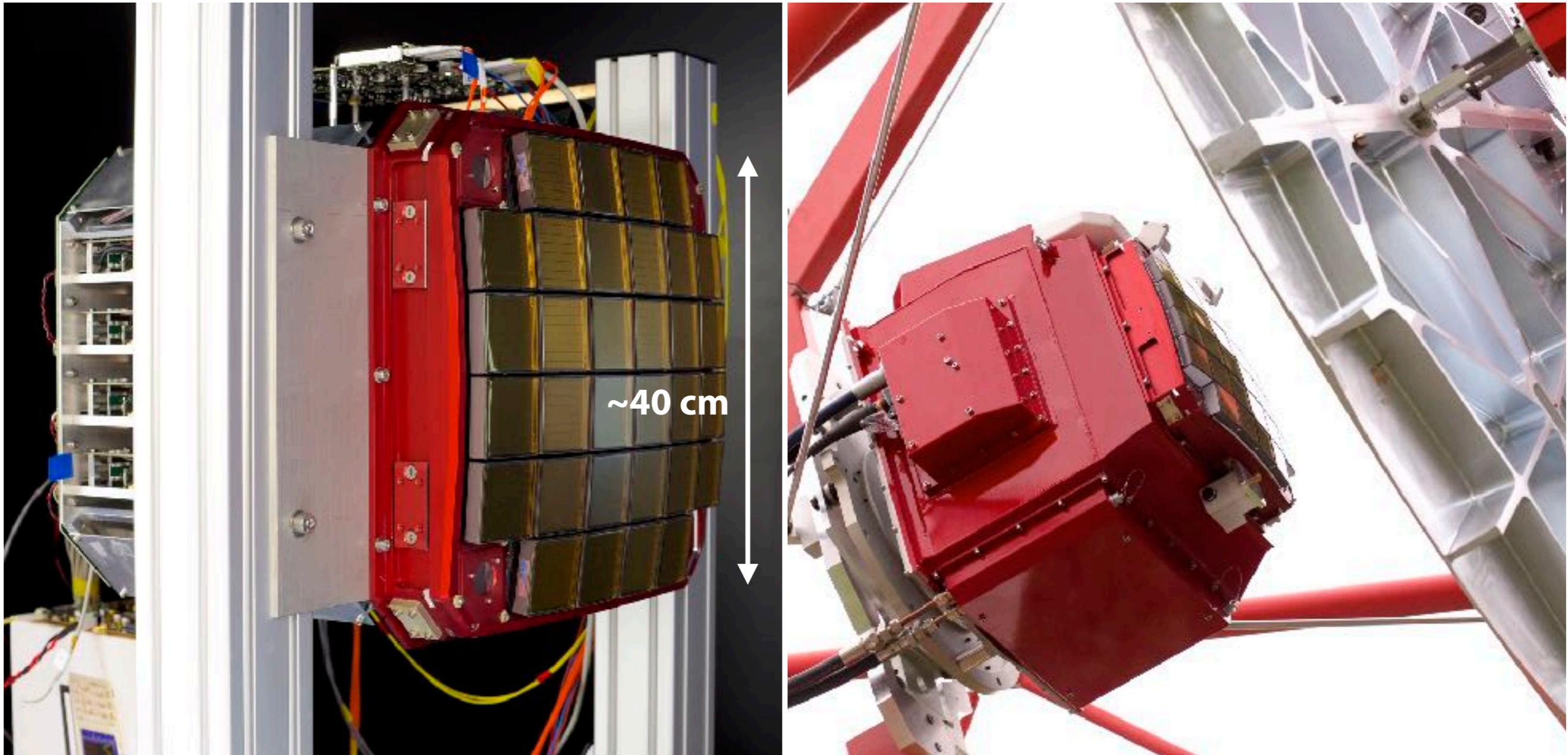
- Optical resolution will be improved ($0.1^\circ \rightarrow < 0.05^\circ$)
- Compact and less expensive camera with small pixels ($\sim 2000 \rightarrow > 10000$ pixels)

CTA Prototypes of Schwarzschild–Couder



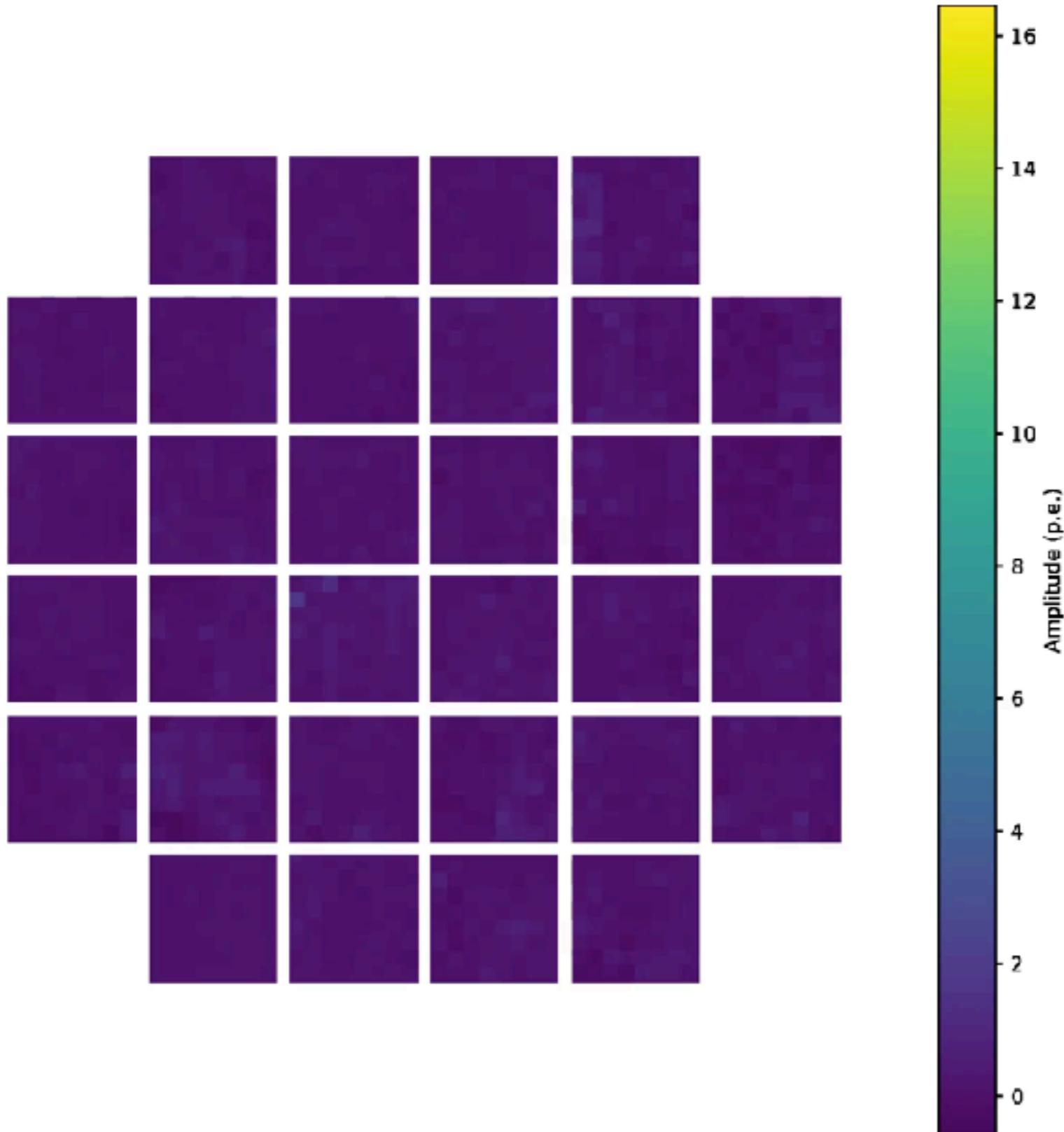
- SC-MST and 2M-SST (GCT) are being developed by ISEE (Nagoya), US, and Europe
- Camera development, optics simulation, and software development by ISEE
- 2M-SST (ASTRI) is also being developed by Italy

Camera Prototype for 2M-SST (GCT)



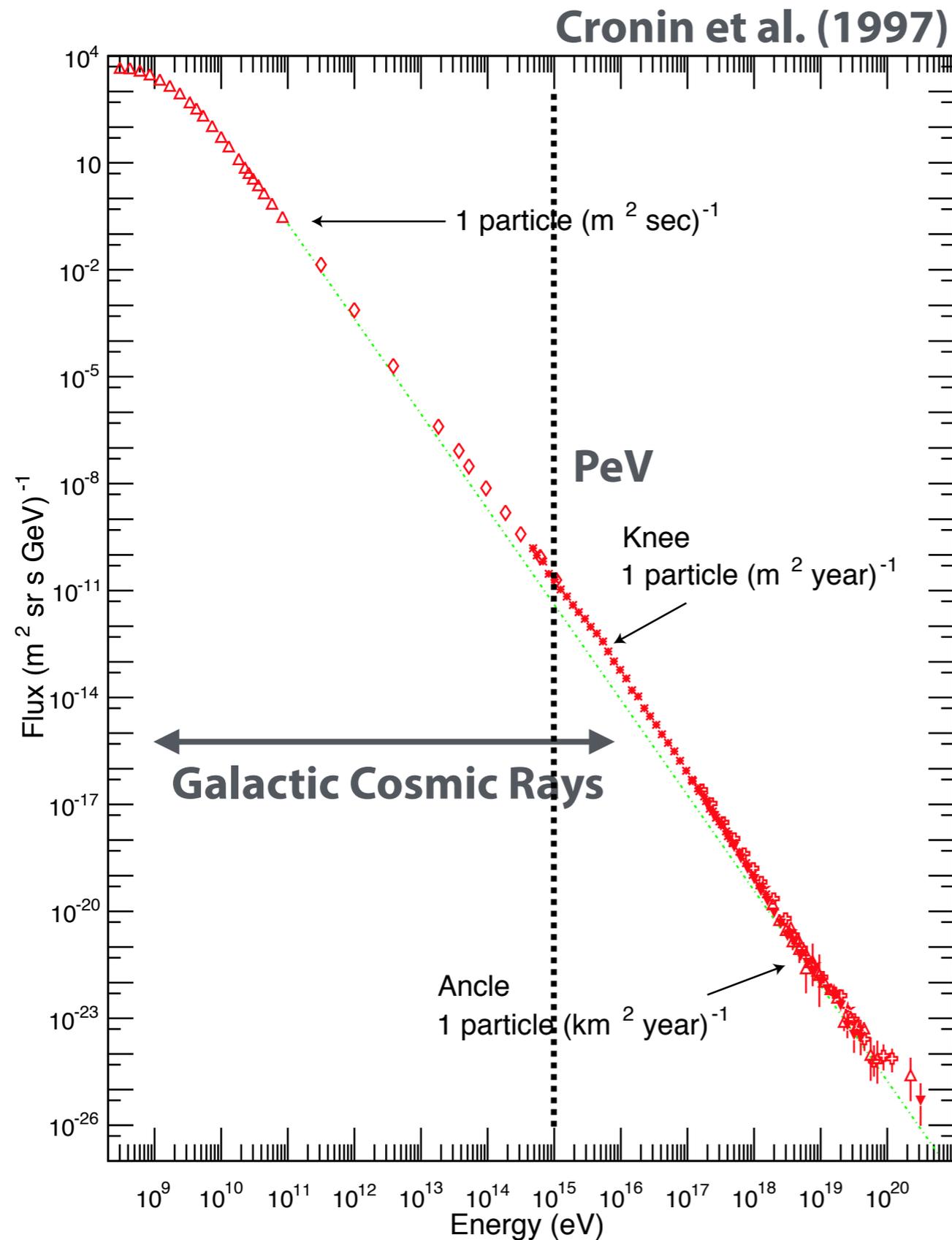
- 2048 pixels with multi-anode PMTs (to be updated to silicon photomultipliers)
- Capable of 1-ns frame “video” recording (i.e., 1 GHz) for Cherenkov flashes (~ 10 ns)
- Installed on the prototype telescope in Nov 2015

Cherenkov Showers



- Achieved the first light at Paris Observatory in 2015
- The first Cherenkov images (proton events) ever in CTA

PeV Cosmic Rays



- Galactic cosmic rays up to ~PeV energies
- Galactic Center and SNRs are leading candidates of **PeVatrons**

CTA Science and the Key Science Projects (KSPs)

- Dark matter

- **KSP: Galactic Center (525 + 300 hours)**

Sgr A* + Halo

- **KSP: Galactic Plane Survey (1020 + 600 hours)**

South + North

- KSP: LMC Survey (340 + 150 hours)

- KSP: Extragalactic Survey

Typical obs. time ~50 hours per object

- KSP: Transients

- **KSP: Cosmic Ray PeVatrons (250 + 50 hours)**

Candidates from GPS + RX J1713

- KSP: Star Forming Systems

- KSP: Active Galactic Nuclei

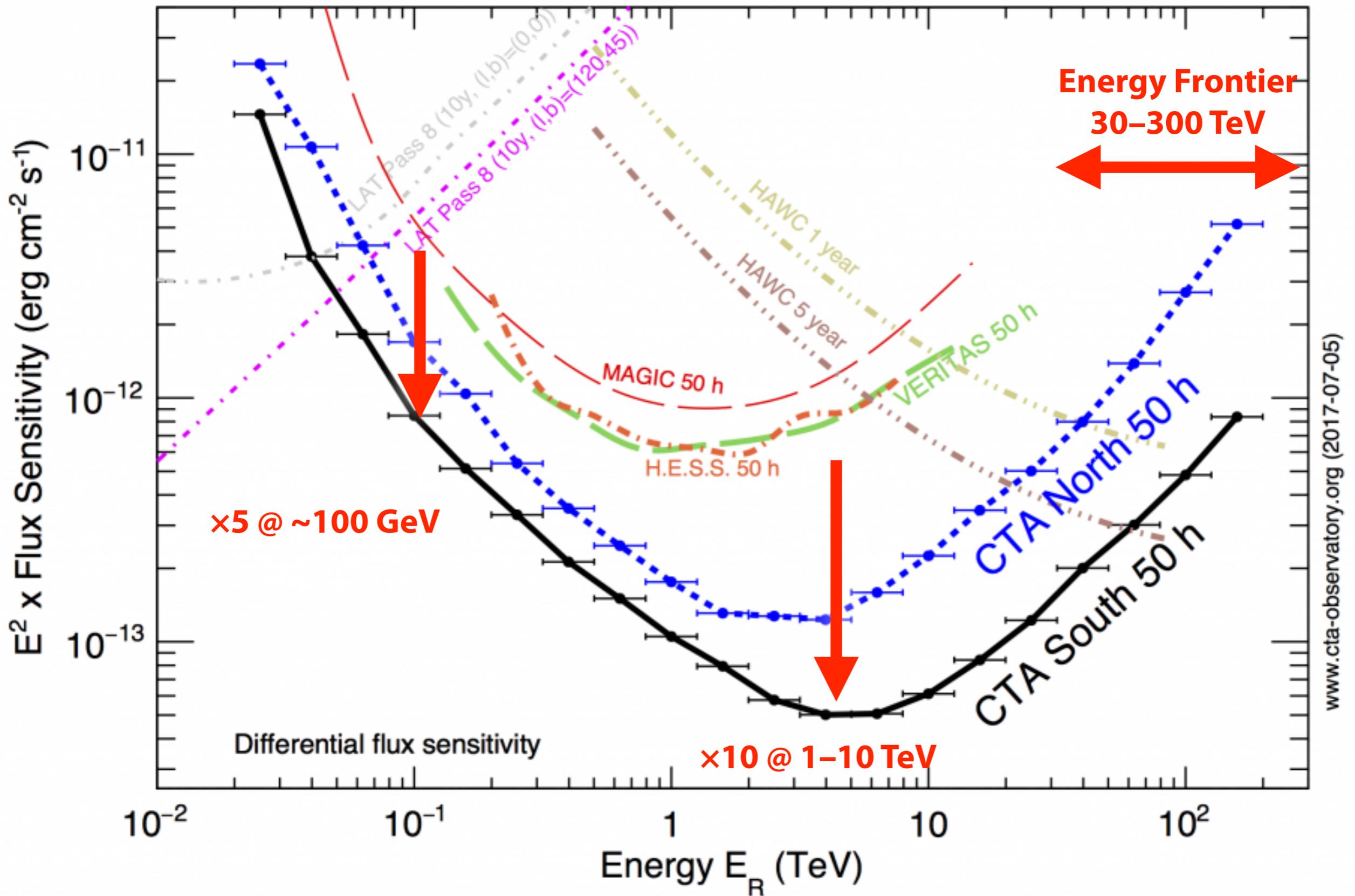
- KSP: Clusters of Galaxies

- Non-Gamma-ray Science

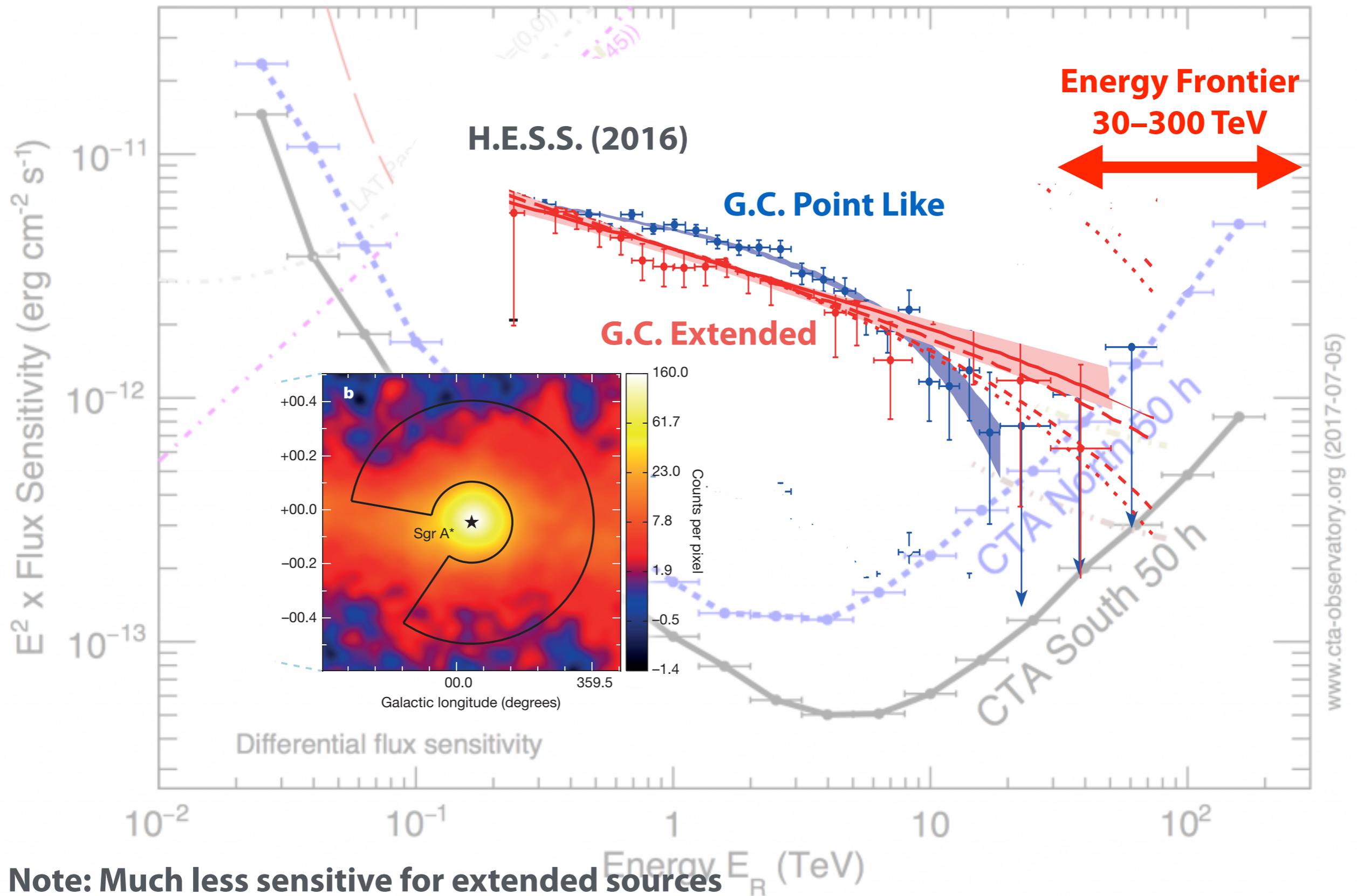
“Science with CTA” will be published soon

Point Source Sensitivity

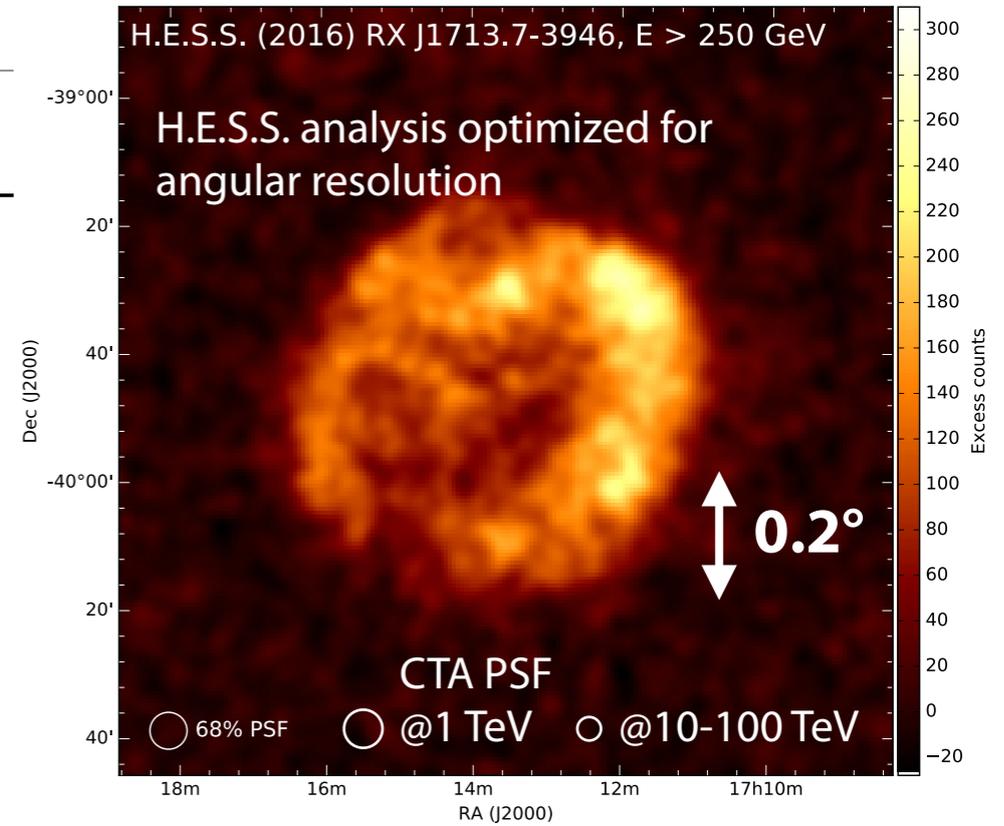
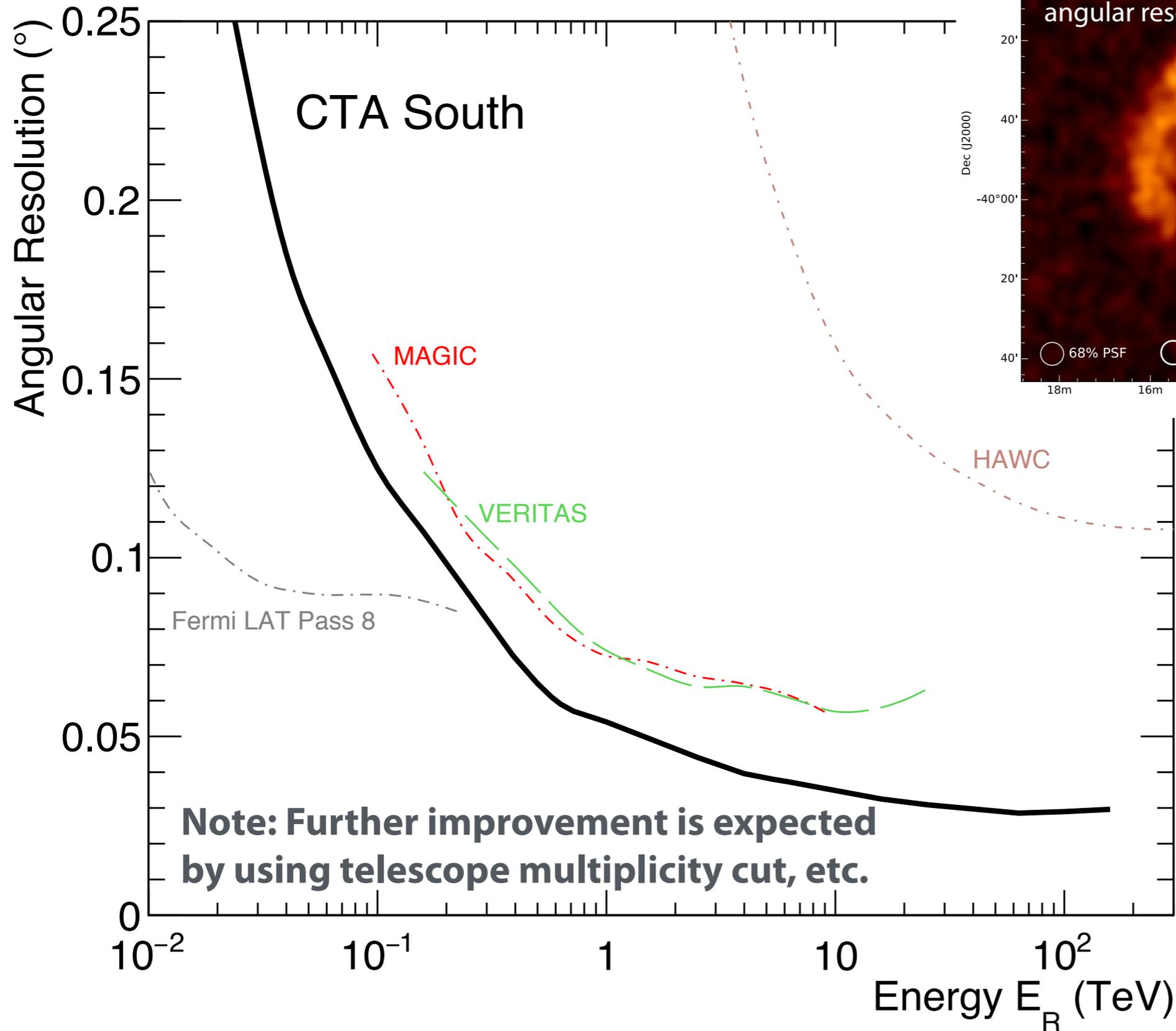
5 bins per decade, point source, high Gal. lat.



Point Source Sensitivity

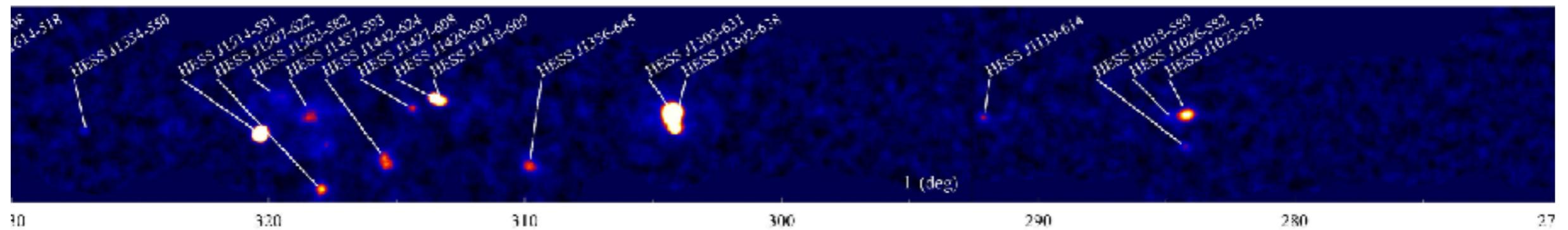
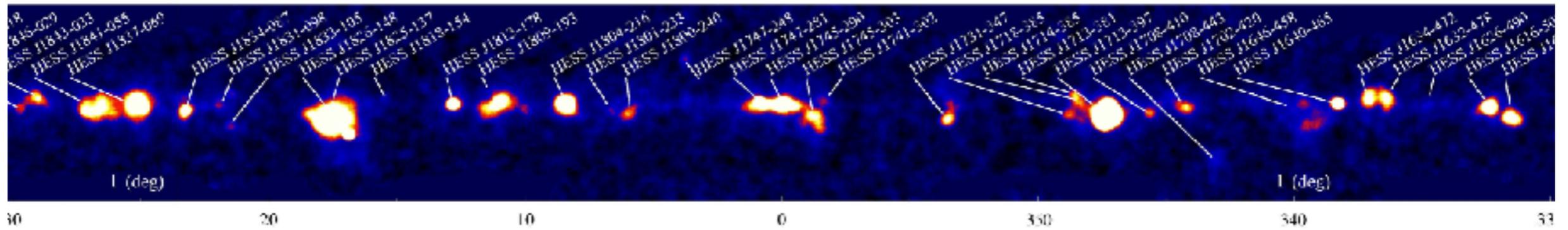
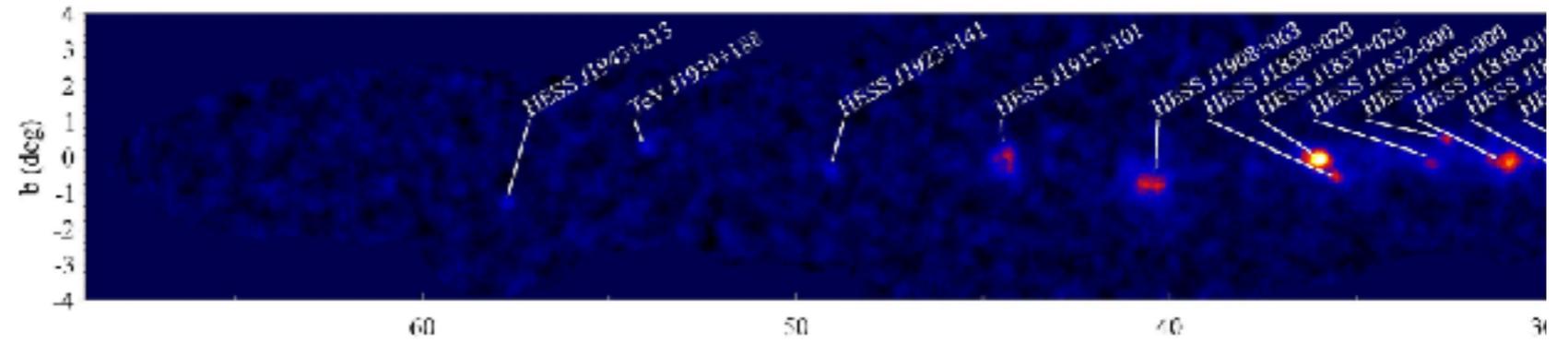


Angular Resolution

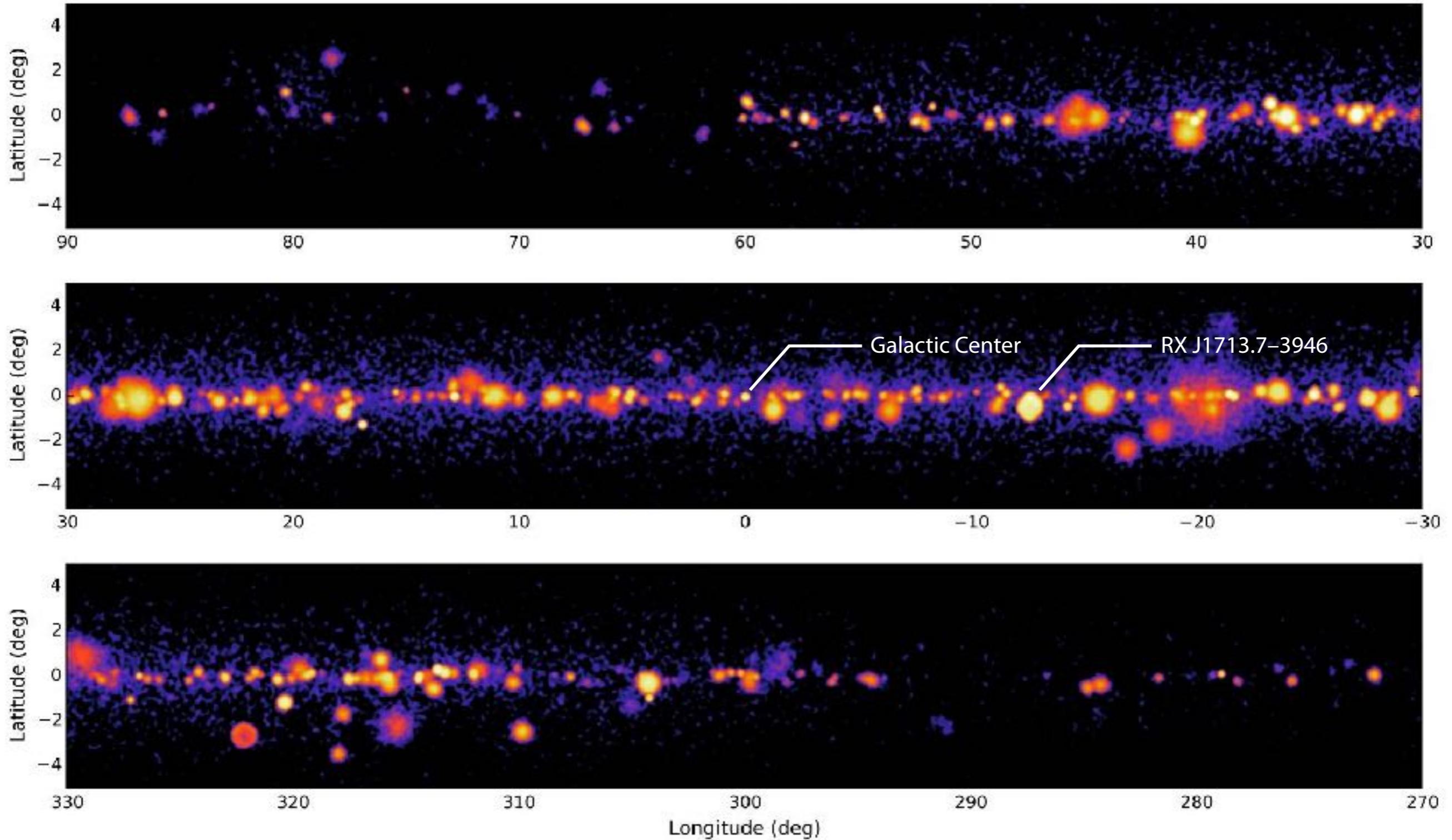


www.cta-observatory.org (2015-05-1)

H.E.S.S. Galactic Plane Survey (as of 2012)

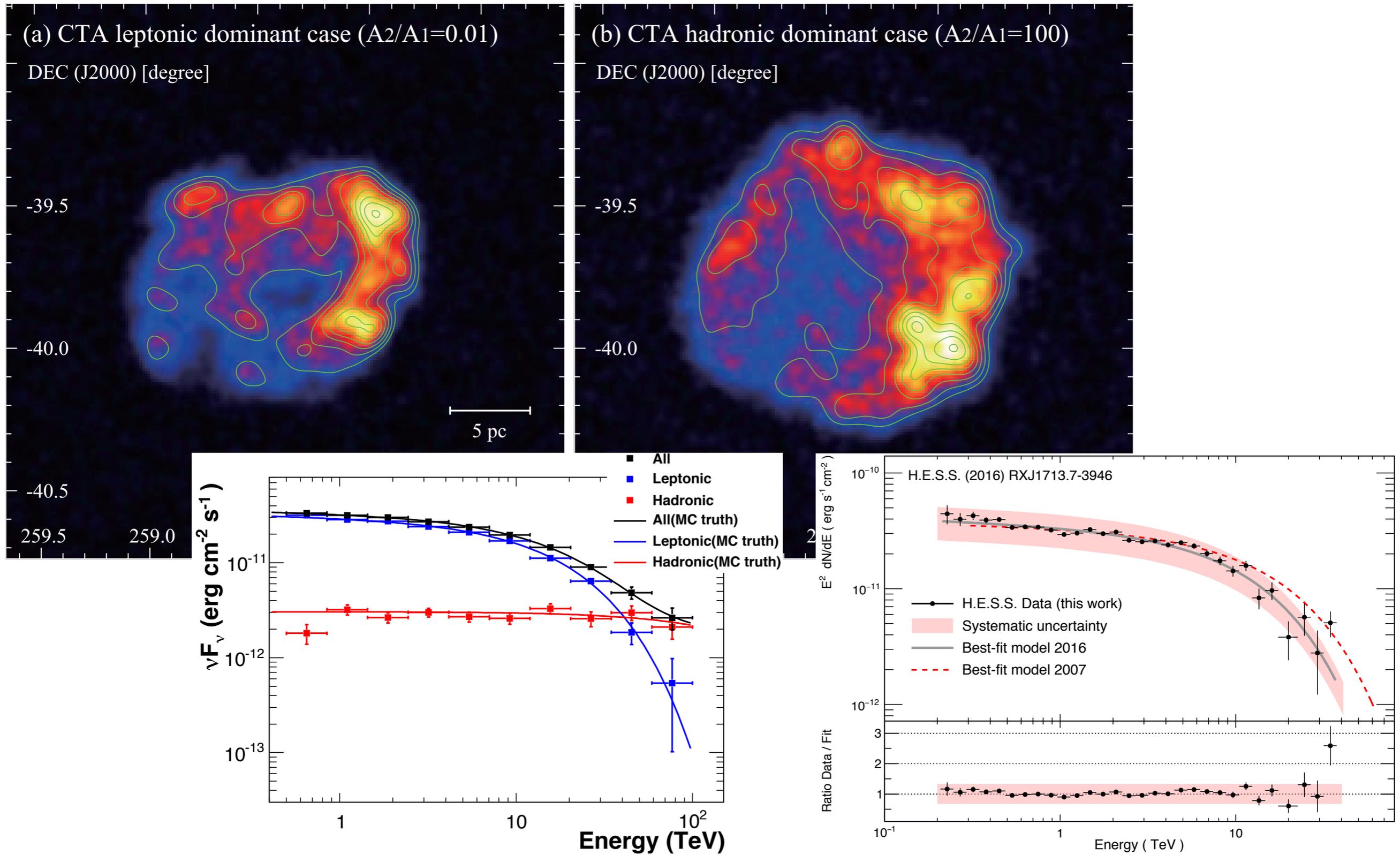


CTA View of the Galactic Plane (Simulation)



CTA View of RX J1713.7–3946 (Simulation)

Nakamori *et al.* (2015)



Summary

- CTA telescopes with Schwarzschild–Couder designs are being prototyped for MSTs and SSTs
- Succeeded in imaging Cherenkov showers for the first time ever in CTA
- Wider FOV, higher angular resolution, and ~100 telescopes will extend the view of very-high-energy sky
 - ▶ Galactic plane survey and PeVatron search
 - ▶ Detailed study of SNRs
 - ▶ Galactic Center